#### Exotic charged current interactions in tritium beta decay

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neutrinos, dark matter & dark energy physics

Talk based on P.O.L., W. Rodejohann (MPIK Heidelberg): arXiv:1603.08690

- I. Direct neutrino mass experiments
- II. Physics of tritium beta decay
- III. Exotic CC interactions in tritium beta decay
- IV. Summary and conclusions

I. Direct neutrino mass experiments

## Neutrino mass bounds from direct neutrino mass experiments

 Bounds from direct neutrino mass experiments: Mainz and Troitsk experiment [hep-ex/0412056; 1108.5034]: Precise study of the endpoint of the tritium beta spectrum.

 $\rightarrow m_{\nu} < 2 \,\mathrm{eV}.$ 

Direct neutrino mass experiments are important to confirm/challenge cosmological observations and  $m_{\beta\beta}$  bounds.

Generic method of direct neutrino mass experiments: Measurement of endpoint of  $\beta$ -spectrum or electron capture spectrum:

$$\mathcal{A} \to \mathcal{B} + e^- + \overline{\nu_e},$$
  
 $\mathcal{A} + e^- \to \mathcal{B} + \nu_e.$ 

#### Direct neutrino mass experiments

Ideal isotopes: Q-value as small as possible  $\rightarrow m_{\nu} \neq 0$  has largest effect on spectrum.

- Tritium:  ${}^{3}\mathrm{H} \rightarrow {}^{3}\mathrm{He}^{+} + e^{-} + \overline{\nu_{e}}$  (Q = 18.6 keV),
- Holmium:  ${}^{163}{
  m Ho}^+ + e^- o {}^{163}{
  m Dy} + \nu_e$  (Q = 2.8 keV)

**Experiments:** See talk by Loredana Gastaldo.

- KATRIN (see talk by Guido Drexlin): tritium beta spectrum,
- Project 8: tritium beta spectrum,
- PTOLEMY (see talk by Alfredo Cocco): tritium beta spectrum,
- NuMECS, HOLMES, ECHo: Holmium electron capture spectrum.

#### What can we learn from direct neutrino mass experiments?

- Study of the endpoint of the energy spectrum: Bounds on the absolute neutrino mass scale,
- $\bullet$  search for eV-scale sterile neutrinos,  $\ldots \rightarrow$  see talk by Guido Drexlin.

#### What else?

Example: The KATRIN tritium source has an activity of about  $10^{11} \ \rm decays$  per second

 $\rightarrow$  If experimental setup can be upgraded for measuring the **whole spectrum**: TRISTAN (see talks by Guido Drexlin and Thierry Lasserre):

 $\rightarrow$  Ultra-high statistics and ultra-high precision measurement of the tritium beta spectrum.

# What can we learn from extensions of KATRIN-like experiments?

Ultra-high statistics and ultra-high precision measurement of the tritium beta spectrum.

This can even lead to bounds on (or discovery of) **new physics** provided that

## the standard model physics involved in beta decay is sufficiently well understood.

For a review on this issue see Susanne Mertens *et al.* [1409.0920] and the talk by Thierry Lasserre.

For the research project presented in this talk: Assumed that sufficient understanding is given.

# What can we learn from extensions of KATRIN-like experiments?

In our paper we studied three effects:

- Difference relativistic / non-relativistic treatment of the spectrum,
- spectral distortion from new charged-current (CC) interactions
- for light active and keV sterile neutrinos

### II. Physics of tritium beta decay

#### Kinematics of $\beta$ decay

$$egin{aligned} \mathcal{A} &
ightarrow \mathcal{B} + e^- + \overline{
u_e} \ |\overline{
u_e}
angle &= \sum_{j=1}^{3+n_s} U_{ej}|
u_j
angle \end{aligned}$$

Fully relativistic treatment of 3-body decay gives

$$\left(\frac{d\Gamma}{dE_e}\right)_{\overline{\nu_j}} = \frac{1}{64\pi^3 m_{\mathcal{A}}} \int_{E_j^-}^{E_j^+} dE_j |\mathcal{M}(\mathcal{A} \to \mathcal{B} + e^- + \overline{\nu_j})|^2$$

- $E_j^{\pm}$  min. and max. neutrino energy (function of particle masses and  $E_e$ ).
- $|\mathcal{M}(\mathcal{A} \to \mathcal{B} + e^- + \overline{\nu_j})|^2$  squared matrix element (unpolarized): function of particle masses and  $E_e$  and  $E_j$ .

#### Theoretical framework

- Lorentz invariance,
- only tree-level interactions,
- only effective four-fermion interactions,
- no assumption about the Lorentz structure  $\to$  all types of interactions allowed (scalar, vector, axial vector, ...)

Under these assumptions the amplitude  $\ensuremath{\mathcal{M}}$  has the form

$$\mathcal{M} = [\overline{u}_e \mathcal{O} v_j] [\overline{u}_{\mathcal{B}} \mathcal{O}' u_{\mathcal{A}}].$$

If operators  ${\cal O}$  and  ${\cal O}'$  independent of particle momenta,  $|{\cal M}|^2$  contains only terms of the form

$$(p \cdot p')$$
 or  $(p \cdot p')(p'' \cdot p''')$ .

#### Theoretical framework

$$\mathcal{M} \sim (p \cdot p')$$
 or  $(p \cdot p')(p'' \cdot p''')$ 

 $\rightarrow$  Energy-momentum conservation:  $p_{\mathcal{A}} = p_{\mathcal{B}} + p_e + p_j$ 

 $\Rightarrow p \cdot p'$  depends only on  $E_e$  (electron energy) and  $E_j$  (neutrino energy).

$$\left|\mathcal{M}(\mathcal{A} \to \mathcal{B} + e^- + \overline{\nu_j})\right|^2 = \mathcal{A} + B_1 E_e + B_2 E_j + C E_e E_j + D_1 E_e^2 + D_2 E_j^2,$$

 $\Rightarrow$  The energy spectrum in our framework can be parameterized by six parameters!

In a given model A,  $B_1$ ,  $B_2$ , C,  $D_1$  and  $D_2$  depend on the particle masses and coupling constants only!

#### Relativistic electron energy spectrum

$$\begin{split} \left(\frac{d\Gamma}{dE_e}\right)_{\overline{\nu_j}} &= \frac{1}{64\pi^3 m_{\mathcal{A}}} \int_{E_j^-}^{E_j^+} dE_j \left|\mathcal{M}(\mathcal{A} \to \mathcal{B} + e^- + \overline{\nu_j})\right|^2 \\ &= \frac{1}{64\pi^3 m_{\mathcal{A}}} \times \\ & \left\{ (\mathcal{A} + \mathcal{B}_1 E_e + \mathcal{D}_1 E_e^2)(E_{j+} - E_{j-}) + \right. \\ & \left. \frac{1}{2} (\mathcal{B}_2 + \mathcal{C} E_e)(E_{j+}^2 - E_{j-}^2) + \right. \\ & \left. \frac{1}{3} \mathcal{D}_2(E_{j+}^3 - E_{j-}^3) \right\}. \end{split}$$

#### Effect of relativistic/non-relativistic spectrum

• Effect on spectral endpoint position:

$$(E_e^{\max})_{\rm NR} = m_A - m_B - m_j,$$
  
 $(E_e^{\max})_{\rm R} = \frac{m_A^2 + m_e^2 - (m_B + m_j)^2}{2m_A}.$ 

- **Difference for tritium decay:**  $\approx$  3.4 eV.
- Whole spectrum: Difference is of the order of  $\approx 10^{-4} \div 10^{-3}.$

In the following: Show *"textbook example":* Standard model expression relativistic/non-relativistic.

#### Example: Standard model: relativistic/non-relativistic

$$\begin{aligned} \mathcal{L} &= -\frac{G_F}{\sqrt{2}} V_{ud} \left( \bar{e} \gamma^{\mu} (\mathbb{1} - \gamma^5) \nu_e \right) \left( \overline{\mathcal{B}} \gamma_{\mu} (g_V \mathbb{1} - g_A \gamma^5) \mathcal{A} \right) + \text{H.c.} \end{aligned}$$

$$\Rightarrow A &= \frac{\gamma}{2} m_{\mathcal{A}} m_{\mathcal{B}} (g_V^2 - g_A^2) (m_{\mathcal{A}}^2 - m_{\mathcal{B}}^2 + m_e^2 + m_j^2), \\ B_1 &= \frac{\gamma}{2} m_{\mathcal{A}} \left\{ (g_V - g_A)^2 (m_{\mathcal{A}}^2 - m_{\mathcal{B}}^2 + m_e^2 - m_j^2) - 2m_{\mathcal{A}} m_{\mathcal{B}} (g_V^2 - g_A^2) \right\}, \\ B_2 &= \frac{\gamma}{2} m_{\mathcal{A}} \left\{ (g_V + g_A)^2 (m_{\mathcal{A}}^2 - m_{\mathcal{B}}^2 - m_e^2 + m_j^2) - 2m_{\mathcal{A}} m_{\mathcal{B}} (g_V^2 - g_A^2) \right\}, \\ C &= 0, \\ D_1 &= -\gamma m_{\mathcal{A}}^2 (g_V - g_A)^2, \\ D_2 &= -\gamma m_{\mathcal{A}}^2 (g_V + g_A)^2, \\ \gamma &\equiv 16 \ G_F^2 |V_{ud}|^2 |U_{ej}|^2. \end{aligned}$$

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#### Example: Standard model: relativistic/non-relativistic

Standard model  $\beta$ -spectrum for tritium decay:



#### Example: Standard model: relativistic/non-relativistic

Non-relativistic approximation is the leading term in the expansion  $\frac{\text{energy scale}}{m_A}$ 

$$\rightarrow \left(\frac{d\Gamma}{dE_e}\right)_{\mathrm{NR}, m_j=0} = \frac{2 G_F^2 |V_{ud}|^2}{\pi^3} |\vec{p}_e| E_e (m_A - m_B - E_e)^2.$$

Deviation relativistic/non-relativistic expression:  $\Delta \equiv \frac{(d\Gamma/dE_e) - (d\Gamma/dE_e)_{\text{NR}}}{(d\Gamma/dE_e)_{\text{NR}}}.$ 



#### Deviation $\sim 10^{-4} \div 10^{-3}$

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III. Exotic CC interactions in tritium beta decay

#### Lagrangian

Effective operator approach: All possible Lorentz-invariant interactions. Use notation of Cirigliano *et al.* [1303.6953].  $L = 1 - \gamma^5$ . Add RH  $\nu$ s.

$$\mathcal{L}_{\rm CC} = -\frac{G_F V_{ud}}{\sqrt{2}} \left\{ (1 + \delta_\beta) (\overline{e} L_\mu \nu_e) (\overline{u} L^\mu d) + \sum_j \overset{(\sim)}{\epsilon_j} (\overline{e} \mathcal{O}_j \nu_e) (\overline{u} \mathcal{O}'_j d) \right\} + \text{H.c.}$$

$$\begin{array}{c} \stackrel{(\sim)}{\epsilon_{j}} & \mathcal{O}_{j} & \mathcal{O}_{j}' \\ \hline \epsilon_{L} & \gamma_{\mu}(\mathbb{1}-\gamma_{5}) & \gamma^{\mu}(\mathbb{1}-\gamma_{5}) \\ \hline \tilde{\epsilon}_{L} & \gamma_{\mu}(\mathbb{1}+\gamma_{5}) & \gamma^{\mu}(\mathbb{1}-\gamma_{5}) \\ \hline \epsilon_{R} & \gamma_{\mu}(\mathbb{1}-\gamma_{5}) & \gamma^{\mu}(\mathbb{1}+\gamma_{5}) \\ \hline \tilde{\epsilon}_{R} & \gamma_{\mu}(\mathbb{1}+\gamma_{5}) & \gamma^{\mu}(\mathbb{1}+\gamma_{5}) \\ \hline \epsilon_{S} & \mathbb{1}-\gamma_{5} & \mathbb{1} \\ \hline \epsilon_{S} & \mathbb{1}+\gamma_{5} & \mathbb{1} \\ \hline -\epsilon_{P} & \mathbb{1}-\gamma_{5} & \gamma^{5} \\ \hline -\tilde{\epsilon}_{P} & \mathbb{1}+\gamma_{5} & \gamma^{5} \\ \hline \epsilon_{T} & \sigma_{\mu\nu}(\mathbb{1}-\gamma_{5}) & \sigma^{\mu\nu}(\mathbb{1}-\gamma_{5}) \\ \hline \tilde{\epsilon}_{T} & \sigma_{\mu\nu}(\mathbb{1}+\gamma_{5}) & \sigma^{\mu\nu}(\mathbb{1}+\gamma_{5}) \end{array}$$

#### Two basic types of couplings: $\epsilon$ : left-handed $\nu$ s, $\tilde{\epsilon}$ : right-handed $\nu$ s. Otranto, 09/09/2016 14 / 21

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#### Numerical analysis

Bounds for the coefficients  $\epsilon$  and  $\tilde{\epsilon}$  from Cirigliano *et al.* [1303.6953].

parameter	best 90 % CL upper bound		used for our estimation
	$ { m Re}\epsilon $	$ \mathrm{Im}\epsilon $	$\epsilon$
$\epsilon_L$	$5 imes 10^{-4}$	$5 imes 10^{-3}$	$5.0 imes10^{-3}$
$\widetilde{\epsilon}_L$	$6 imes 10^{-2}$		$8.5 imes10^{-2}$
$\epsilon_R$	$5 imes 10^{-4}$	$5 imes 10^{-4}$	$7.1 imes10^{-4}$
$\widetilde{\epsilon}_R$	$5 imes 10^{-3}$	$5 imes 10^{-3}$	$7.1 imes10^{-3}$
$\epsilon_{S}$	$8 imes 10^{-3}$	$1 imes 10^{-2}$	$1.3 imes10^{-2}$
$\widetilde{\epsilon}_{S}$	$1.3 imes10^{-2}$	$1.3 imes10^{-2}$	$1.8 imes10^{-2}$
$\epsilon_P$	$4 imes 10^{-4}$	$2 imes 10^{-4}$	$4.5 imes10^{-4}$
$\tilde{\epsilon}_P$	$2  imes 10^{-4}$	$2 imes 10^{-4}$	$2.8 imes10^{-4}$
$\epsilon_T$	$1 imes 10^{-3}$	$1 imes 10^{-3}$	$1.4 imes10^{-3}$
$\widetilde{\epsilon}_T$	$3  imes 10^{-3}$	$3 imes 10^{-3}$	$4.2 imes10^{-3}$

#### Procedure for numerical analysis

- Fixed neutrino masses  $m_j = 0.5 \, \mathrm{eV}$  (light active neutrinos) or
- $m_j = 5 \text{ keV}$  (heavy sterile right-handed neutrino), see talk by Alex Merle,
- one  $\epsilon$  or  $\widetilde{\epsilon}$  set to 1, the others set to 0,
- relevant mixing matrix element set to 1.
- $\rightarrow$  Computed the values for the six coefficients  $A, \ldots, D_2$ .
- $\rightarrow$  Tables in paper.

Paper gives prescription how to scale the tabulated values for other values of  $\epsilon$  and other mixing matrix elements.

## $\Rightarrow$ Can directly obtain the beta spectrum for any new CC interaction.

#### Signal to be expected in a high-precision experiment

Mixing active/sterile (left/right) is of the order of

$$rac{m_{
m light}}{m_{
m heavy}}\lesssim rac{{
m eV}}{{
m keV}}\sim 10^{-3}.$$

 $\Rightarrow$  "heavy neutrinos/ $\epsilon$ " and "light neutrinos/ $\epsilon$ ": strongly suppressed by mixing matrix  $\rightarrow$  no observable effect.

- $\epsilon$ : studied light neutrinos ( $m_j = 0.5 \, \text{eV}$ ),
- $\tilde{\epsilon}$ : heavy neutrinos ( $m_j = 5 \text{ keV}$ ).
- Used upper bounds for  $\epsilon$  and  $\widetilde{\epsilon}$  from literature.

 $\rightarrow$  Plots of:

$$\Delta({\stackrel{(\sim)}{\epsilon_{j}}}) \equiv \frac{\text{test spectrum (NP)} - \text{reference spectrum (no NP)}}{\text{reference spectrum (no NP)}}$$

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#### New physics effects for light neutrinos with $m_j = 0.5 \, \mathrm{eV}$



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#### New physics effects for heavy neutrinos with $m_i = 5 \text{ keV}$



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#### Summary

- Upgraded KATRIN (or KATRIN-like) experiment may have access to the full energy spectrum of tritium decay. → Ultra-high statistics and ultra-high precision!
- Fully relativistic calculation of the  $\beta$  spectrum. Generic spectrum depends on six process-dependent parameters.
- Departure from usual non-relativistic approximation:  $\approx 10^{-4} \div 10^{-3}$ .
- New physics in  $\beta$  decay: Studied all possible new CC interactions in effective operator framework.
- Both light (sub-eV) and heavy (keV) neutrinos considered: Effect on endpoints negligible, full spectrum can show sizable distortions at the permille level.

#### Conclusions

- Accessibility of the new-physics effects by a future KATRIN-like experiment: Example: keV scale right-handed neutrinos: Sensitivity estimate (Mertens *et al.* [1409.0920]):  $\sim 10^{-7}$ .
- $\bullet \rightarrow$  modified KATRIN-like setup sensitive to
  - *ε<sub>L</sub>*, *ε<sub>R</sub>*, *ε<sub>S</sub>*, *ε<sub>T</sub>* in case of light left-handed (even almost massless) neutrinos.

Different new-physics scenarios can be distinguished by the shape of the spectral distortion;

- *ϵ̃<sub>L</sub>*, *ϵ̃<sub>F</sub>*, *ϵ̃<sub>S</sub>*, *ϵ̃<sub>T</sub>* in case of keV-scale right-handed neutrinos.
   New physics effects are not easily distinguishable (shapes are quite similar).
- If systematic effects in the experiment and all Standard Model contributions are under control, an extended KATRIN-like setup may significantly improve the bounds on new CC interactions in β decay.

### Thank you for your attention!



## Backup slides

#### Current bounds on neutrino masses

• Bounds from cosmology: Bounds on the sum of the three active neutrino masses:

$$\begin{split} &\sum m_{\nu} < 0.72 \text{ eV} \quad Planck \text{ TT+lowP}, \\ &\sum m_{\nu} < 0.21 \text{ eV} \quad Planck \text{ TT+lowP+BAO}, \\ &\sum m_{\nu} < 0.49 \text{ eV} \quad Planck \text{ TT}, \text{TE}, \text{EE+lowP}, \\ &\sum m_{\nu} < 0.17 \text{ eV} \quad Planck \text{ TT}, \text{TE}, \text{EE+lowP+BAO}. \end{split}$$



(Planck Collaboration [1502.01589]).

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Dependent on which data taken into account. In any case bound stronger than current direct neutrino mass bounds.

• Bounds from  $(\beta\beta)_{0\nu}$ -searches: Bounds on

$$m_{\beta\beta} = \left|\sum_{k=1}^{3} U_{ek}^2 m_k\right| \rightarrow \text{bounds on absolute mass scale}$$

#### Current bounds on neutrino masses

Dell'Oro et al. [1601.07512]



 $m_
u \lesssim \mathcal{O}(1\,\mathrm{eV})$ 

#### Minimal and maximal neutrino energy

$$E_{j\pm} = \frac{-(m_{\mathcal{A}} - E_{e})(E_{e}m_{\mathcal{A}} - \alpha) \pm |\vec{p}_{e}| \sqrt{(E_{e}m_{\mathcal{A}} - \alpha + m_{j}^{2})^{2} - m_{\mathcal{B}}^{2}m_{j}^{2}}}{m_{\mathcal{A}}^{2} - 2m_{\mathcal{A}}E_{e} + m_{e}^{2}}$$
$$\alpha = \frac{1}{2} \left( m_{\mathcal{A}}^{2} - m_{\mathcal{B}}^{2} + m_{e}^{2} + m_{j}^{2} \right).$$

#### Kurie plots





 $m_i = 2.0, 1.0, 0.5 \text{ and } 0 \text{ eV},$  10, 5, 3 and 0 keV.

#### Momentum-dependent operators $\mathcal{O}$ , $\mathcal{O}'$

In our study this case appears only for the weak magnetism correction:

$$\langle p(p_p) | \overline{u} \gamma_{\mu} d | n(p_n) \rangle = \overline{u}_p(p_p) \left[ g_V(q^2) \gamma_{\mu} - i \frac{g_{WM}(q^2)}{2M_N} \sigma_{\mu\nu} q^{\nu} \right] u_n(p_n) + \mathcal{O}((q/M_N)^2)$$

$$q = p_{\mathcal{A}} - p_{\mathcal{B}}$$

Gives contribution to  $|\mathcal{M}|^2$  of the form

$$\frac{(p \cdot q)(p' \cdot q)(p'' \cdot p''')}{M_N^2} \quad \text{or} \quad \frac{(p \cdot p')(p'' \cdot p''')(q \cdot q)}{M_N^2}.$$

For tritium decay suppressed by

$$rac{q^2}{M_{\mathcal{A}}^2} \lesssim 10^{-10}.$$

#### Corrections from Standard Model physics

See Susanne Mertens et al. [1409.0920] and the talk by Thierry Lasserre:

- Excited final states: the effect on the spectrum is very large—larger than 10 % close to the endpoint. Far from the endpoint  $\sim$  1%.
- Coulomb interaction between the outgoing electron, the daughter nucleus ( $\rightarrow$  Fermi function  $F(Z, E_e)$ ) and the left behind orbital electron of the former <sup>3</sup>H<sub>2</sub>-molecule.
- The nuclear recoil: automatically taken into account by using the exact relativistic expression for  $d\Gamma/dE_e$ .
- The daughter nucleus  $^{3}\mathrm{He}^{2+}$  is not pointlike  $\rightarrow$  modifies the Coulomb field acting on the emitted electron.
- Radiative corrections: The dominant radiative corrections will be QED-corrections.  $\rightarrow$  Of the order of  $\sim 1$  %.
- Hadronic matrix elements.

Moreover, source is a gas at finite temperature (30K).

#### Corrections from Standard Model physics

Hadronic matrix elements: See Cirigliano et al. [1303.6953].

$$\begin{split} \langle p(p_p) | \overline{u} \gamma_{\mu} d | n(p_n) \rangle &= \overline{u}_p(p_p) \left[ g_V(q^2) \gamma_{\mu} - i \frac{g_{WM}(q^2)}{2M_N} \sigma_{\mu\nu} q^{\nu} \right] u_n(p_n) + \mathcal{O}((q/M_N)^2) \\ \langle p(p_p) | \overline{u} \gamma_{\mu} \gamma_5 d | n(p_n) \rangle &= g_A(q^2) \, \overline{u}_p(p_p) \gamma_{\mu} \gamma_5 u_n(p_n) + \mathcal{O}((q/M_N)^2), \\ \langle p(p_p) | \overline{u} d | n(p_n) \rangle &= g_S(q^2) \, \overline{u}_p(p_p) u_n(p_n), \\ \langle p(p_p) | \overline{u} \gamma_5 d | n(p_n) \rangle &= g_P(q^2) \, \overline{u}_p(p_p) \gamma_5 u_n(p_n) = \mathcal{O}(q/M_N), \\ \langle p(p_p) | \overline{u} \sigma_{\mu\nu} d | n(p_n) \rangle &= g_T(q^2) \, \overline{u}_p(p_p) \sigma_{\mu\nu} u_n(p_n) + \mathcal{O}(q/M_N). \end{split}$$

Written here for proton and neutron. The same structure for tritium/ $^{3}$ He.

Values for form factors from measurements or lattice.

#### Neutrino mass states

$$\begin{pmatrix} U & S \\ T & V \end{pmatrix}^{\dagger} M_{\nu} \begin{pmatrix} U & S \\ T & V \end{pmatrix}^{*} = \operatorname{diag}(m_{1}, m_{2}, m_{3}, M_{1}, \dots, M_{s})$$

$$\nu_{L} = U \nu'_{L} + S N'_{R}{}^{c},$$
  
$$\nu_{R} = T^{*} \nu'_{L}{}^{c} + V^{*} N'_{R}.$$

Choose (for numerical estimates) neutrino mass spectrum: light neutrinos with  $m_j = 0.5 \text{ eV}$  and heavy neutrinos with  $m_j = 5 \text{ keV}$  $\rightarrow$  see talk by Alex Merle.