

STOCHASTIC MODEL FOR THE MOTION OF CORRELATED PHOTON PAIRS

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Received 16 September 1981

Revised manuscript received 19 January 1982

If one analyzes the stochastic behaviour of two massive ($m_\gamma \neq 0$) photons imbedded in Dirac's vacuum one obtains (with stochastic jumps at velocity of light) the two-particle Proca equations which one can use to interpret (i) the first results of the Aspect experiment; (ii) the future issues of the complete Aspect and Rapisarda experiments, if they will violate Bell's inequality.

The aim of the present letter is to extend to the two-photon case the relativistic stochastic treatment used by Cufaro Petroni and Vigier [1] and Namsrai [2] in the two-scalar-particle case. As one knows this treatment rests upon

(i) the generalization to an eight-dimensional configuration space of the stochastic formalism given by Guerra and Ruggiero [3];

(ii) a perfectly causal action-at-a-distance description of the relativistic stochastic quantum potential which appears in the said formalism [4,5]. This is necessary, if one wants to establish a realistic causal explanation [6] of quantum nonseparability. No experiments have yet been performed to test whether the latter is a real phenomenon, but a series of related experiments, culminating in the recent one of Aspect et al. [7], suggests to us that, in this as in all previous tests, the predictions of quantum theory will be confirmed, and, in particular, that they will be confirmed in the photon-correlation experiments proposed by Aspect [8] and Rapisarda [9]. Of course, if the experiments, instead of confirming the phenomenon of quantum nonseparability, should turn out to give a result consistent with Bell's inequality, then it would be necessary to examine "modifications" of quantum

theory rather than new "interpretations", because of the failure of a key prediction of that theory.

Our demonstration is based upon the introduction of an eight-dimensional configuration space where a pair position is defined by an eight component vector X^i ($i = 1, \dots, 8$) where:

$$\{X^i\}_{i=1,\dots,8} \equiv \{x_1^\mu; x_2^\nu\}_{\mu,\nu=0,\dots,3}, \quad (1)$$

with x_1^μ, x_2^ν four-vectors of the position of each body [1,2], so that

$$X^2 = X_i X^i = (x_1)^2 + (x_2)^2. \quad (2)$$

If $x_1^\mu(\tau_1), x_2^\nu(\tau_2)$ are the trajectories for the two photons, the trajectory in the configuration space will be $X^i(\tau_1, \tau_2)$. As a consequence of Nelson's equations [10] we can now generalize the differential operators defined by Guerra and Ruggiero [3] (which we have already used to demonstrate Proca's and Dirac's equations for isolated particles [11]) to a system of two photons. We thus write:

$$\begin{aligned} D &= \partial/\partial\tau_1 + \partial/\partial\tau_2 + b_i \partial^i, \\ \delta D &= -(2m_\gamma)^{-1} \square + \delta b^i \partial_i, \quad \partial^i = \partial/\partial X_i, \\ \square &= \partial_i \partial^i = \square_1 + \square_2, \quad b_i = DX_i, \quad \delta b_i = \delta DX_i \end{aligned} \quad (3)$$

(with $\hbar = c = 1$).

We now suppose (as we have done for the Feynman and Gell-Mann equation [11]) that the drift and stochastic velocities b_i and δb_i can be derived from:

$$b^i = \partial^i \Phi / m_\gamma, \quad \delta b^i = -\partial^i \ln \rho^2 / 2m_\gamma = -m_\gamma^{-1} (\partial^i \rho) / \rho, \quad (4)$$

where $\Phi(X^i, \tau_1, \tau_2)$ and $\rho(X^i)$ are real tensors with two indices, which can be represented by 4×4 matrices, the indices μ, ν taking the values $0, \dots, 3$. We also require that ρ is nonsingular, that ρ^2 is a trace-invariant matrix (which plays the role of a density matrix), and finally that Φ and ρ are commuting matrices.

The classical covariant stochastic equation and the continuity equations for our photon pair are

$$(DD - \delta D \delta D) X^i = 0, \quad \partial \rho / \partial \tau_1 = -\partial_{1\mu} (\rho b_1^\mu), \quad \partial \rho / \partial \tau_2 = -\partial_{2\nu} (\rho b_2^\nu). \quad (5)$$

Of course these equations are two-indices (4×4) tensor equations. In our model, as a generalization of the assumption [3] that ρ is independent of the proper time in the one-body case, we make the physical hypothesis that the total number of particles is conserved, and thus we write

$$\partial \rho / \partial \tau_1 + \partial \rho / \partial \tau_2 = 0, \quad (6)$$

so that our continuity equation in the configuration space is

$$\partial_i (b^i \rho) = 0. \quad (7)$$

From eqs. (3) and (4) we have for the stochastic covariant equation (5)

$$\partial_i \left(\frac{1}{2m_\gamma} \frac{\square \rho}{\rho} - \frac{\partial \Phi}{\partial \tau_1} - \frac{\partial \Phi}{\partial \tau_2} - \frac{1}{2m_\gamma} \partial_j \Phi \partial^j \Phi \right) = 0, \quad (8)$$

which shows that the term in the parentheses is constant. If we pose this constant equal to zero, by rearranging the energy scale, eq. (8) takes the following form

$$\left(\square - 2m_\gamma \frac{\partial \Phi}{\partial \tau_1} - 2m_\gamma \frac{\partial \Phi}{\partial \tau_2} - \partial_j \Phi \partial^j \Phi \right) \rho = 0. \quad (9)$$

Similarly for the continuity equation (5) we have from eqs. (3) and (4)

$$(2\partial_j \Phi \partial^j + \square \Phi) \rho = 0. \quad (10)$$

If we now make a linear combination of eqs. (9) and (10) with coefficient 1 and i , we have after multiplication by $\exp(i\Phi)$:

$$\left[\left(\square + 2i\partial^j \Phi \partial_j + i\square \Phi - \partial_j \Phi \partial^j \Phi - 2m_\gamma \frac{\partial \Phi}{\partial \tau_1} - 2m_\gamma \frac{\partial \Phi}{\partial \tau_2} \right) \rho \right] \exp(i\Phi) = 0, \quad (11)$$

which takes the following synthetic form

$$\left[\square + 2im_\gamma \frac{\partial}{\partial \tau_1} + 2im_\gamma \frac{\partial}{\partial \tau_2} \right] \rho \exp(i\Phi) = 0. \quad (12)$$

With the notations

$$\Phi(X^j; \tau_1, \tau_2) = S(X^j) - m_\gamma(\tau_1 + \tau_2); \quad A = \rho \exp(i\Phi), \quad (13)$$

we have finally (A is a two-indices tensor)

$$[\square + 2m_\gamma^2] A = 0. \quad (14)$$

This is the wave equation that generalizes Proca's equation to correlated photon pairs represented (as suggested by Landau and Lifshitz [12]) by a tensor field formalism. As Landau and Lifshitz have shown, the singlet corresponds to the antisymmetric part of the $A_{\mu\nu}$ field, the triplet is described by the symmetric component.

We conclude with two remarks. The first is that the preceding calculations represent real physical collective motions on the top of a covariant stochastic "vacuum" (i.e. Dirac's aether [13]). The pair's motion along correlated world lines in real space-time are represented by single world lines in configuration space. The nature of their correlations results from the model since the presence of the particle 1 disturbs the aether i.e. the motion of the particle 2 and vice versa. Physically this amounts (in the hydrodynamical model of Bohm and Vigier [14]) to the superposition in the space-time of two interacting fluids which undergo lightlike internal stochastic motions, particle-antiparticle transitions and possible number-preserving transfers from one fluid to another ... so that we have a conserved tensor fluid particle density in configuration space.

The second remark is that eq. (14) extends the quantum potential action-at-a-distance [5] to spin-spin correlations as will be shown in a subsequent letter. These actions-at-a-distance are carried by intervening

stochastic extended vacuum particles which transmits nonlocal interactions in their interior [13].

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