

Luminosity Distance and Non Perturbative Cosmology

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Importance of an exact non-perturbative approach

- ▶ Theoretical motivations:
 - Considering all of the effects due to inhomogeneities and anisotropies
 - Evaluating the impact of the backreaction
 - Non linearity of Einstein equations
- ▶ Experimental reasons:
 - Comparing the prediction with respect to the next generation measurements of the precision cosmology era
 - Re-considering the right averaging procedure for the physical observables

The Geodesic Light-Cone gauge

- ▶ New approach to the study of cosmology
- ▶ Fully non perturbative
- ▶ Gauge fixing of the $g_{\mu\nu}$ on the observer

$$ds^2 = \Upsilon^2 dw^2 - 2\Upsilon dw d\tau + \gamma_{ab} \left(d\tilde{\theta}^a - U^a dw \right) \left(d\tilde{\theta}^b - U^b dw \right)$$

- ▶ Light-cone condition: $w = \text{constant}$
- ▶ Geodesic flow: $u_\mu = -\partial_\mu \tau$
- ▶ Ability to average on the Light-cone
(Gasperini, Marozzi, Nugier, Veneziano, JCAP 1107
(2011) 008)

Luminosity distance in this gauge (1)

- ▶ Exact expression for some observable quantities as redshift
- ▶ Exact non perturbative expression for the Jacobi map (Fanizza, Gasperini, Marozzi, Veneziano, JCAP 1311 (2013) 019)
- ▶ This allows us to get an exact expression for the luminosity distance:

$$d_L^2 = \left(\frac{\gamma_o}{\gamma_s} \right)^4 \frac{4\sqrt{\gamma_s}}{[\det(u_\tau^{-1} \partial_\tau \gamma^{ab}) \gamma^{3/2t}]_o}$$

Luminosity distance in this gauge (2)

- ▶ First exact solution of the Jacobi equation
- ▶ The expression appears as a simple product of an observer term times a source one
- ▶ True only in this gauge
- ▶ The result takes care of the right physical phenomena at the observer position, due to his peculiar velocity
- ▶ Transformation into the Longitudinal gauge, the Synchronous gauge and among them is fully consistent

Future perspectives

- ▶ Solving Einstein equations in this gauge
- ▶ Expressing other physical observables with the aim of this gauge (CMB power spectrum, Weak lensing...)
- ▶ Averaging over partial region of the sky, in order to compare the local quantities with global ones