



# Neutrino Mass Models in the LHC Era



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# Outline of Talk

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0. Seesaw mechanism
1. Seesaw Scale: Theory insights, LHC, other searches
2. Large mixings and unification with quarks
3. A new possibility for neutrino masses

(only three active neutrinos will be considered-MINOS and MiniBooNe not discussed; see E. Akhmedov's talk)

# Neutrino Mass

→ New physics

■ Standard model- no  $\nu_R \rightarrow m_\nu = 0$

■ Two possibilities for  $m_\nu$

(i) SM +  $\nu_R \rightarrow L_Y = h_\nu \bar{L} H \nu_R + h.c. \rightarrow m_\nu = h_\nu v_{wk}$

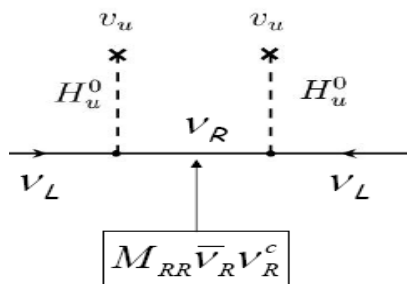
Observations →  $h_\nu \cong 10^{-12}$  : any justification ?

(ii) Mass from high scale physics:  $\frac{LHLH}{M}$

# What is the new physics ?

- Seesaw paradigm:** New fermions  $\mathbf{N}_R$  or Higgs

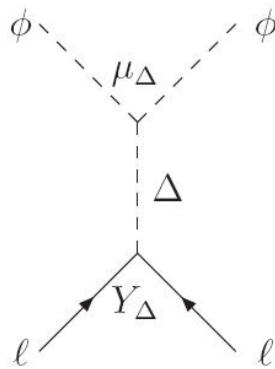
**Type I**



$$m_\nu \cong -\frac{h_\nu^2 v_{wk}^2}{M_R}$$

**N**- Majorana

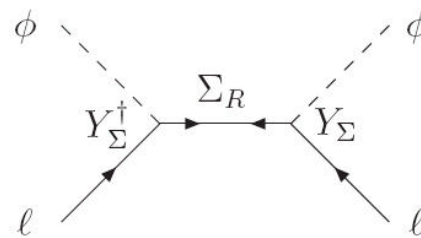
**Type II**



$$m_\nu = -2Y_\Delta v^2 \frac{\mu_\Delta}{M_\Delta^2}$$

Triplet Higgs

**Type III**



$$m_\nu = -\frac{v^2}{2} Y_\Sigma^T \frac{1}{M_\Sigma} Y_\Sigma$$

Triplet Fermion

**Inverse**

$$\begin{pmatrix} 0 & h v_{wk} & 0 \\ h v_{wk} & 0 & M \\ 0 & M & \mu \end{pmatrix}$$

$$m_\nu \cong -m_D^T M^{-1} \mu M^{-1} m_D$$

**N** - Dirac



# Seesaw scale ?

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- **Seesaw assumption by itself cannot tell –need to know Dirac mass**
- **Plausible guesses motivate different scales**

# Is seesaw physics accessible at LHC ?

- **Type I case, if  $m_D \approx m_e$  scale  $\sim$  TeV**

- **Inverse seesaw**  $\rightarrow$  scale at TeV even with  $m_D \approx m_t$

- if there is a 4<sup>th</sup> gen., M necessarily TeV.

**Scale being TeV- no guarantee of LHC visibility:**

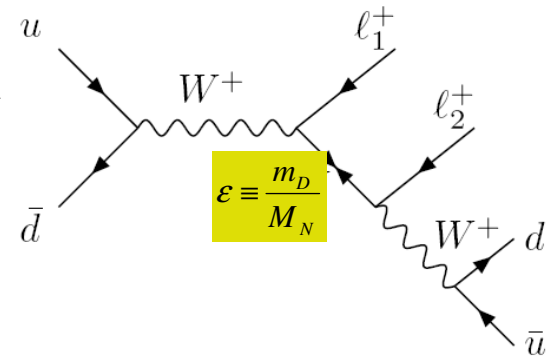
**key parameter:**  $\varepsilon \equiv \frac{m_D}{M_N} < 10^{-7}$  (type I);  $> 10^{-3}$  (Inverse)

# LHC signal for simplest TeV seesaw

- Type I and Inverse:
- Basic manifestations of seesaw:  $\nu - N$  mixing

**Collider production** for

$$M_N < TeV$$



- **Negligible for type I seesaw but observable for inverse seesaw  $M_N \sim TeV$**
- **Situation different with other seesaws and seesaw with gauge forces !!**

# Type II and III at LHC

- New particles are SM non-singlets:

- **Type II**:  $\Delta = \begin{pmatrix} \delta^+/\sqrt{2} & \delta^{++} \\ \delta^0 & -\delta^+/\sqrt{2} \end{pmatrix}$ . **couples only to leptons**

$$u\bar{u} \rightarrow \delta^{++} \delta^{--}; u\bar{d} \rightarrow \delta^{++} \delta^-$$

**directly probes neutrino mass matrix.**

**4-lepton signal**  $l^+ l^+ l^- l^-$  (Han, Perez, Huang, Li, Wang;...)

- **Type III**:  $\Sigma = \begin{pmatrix} \Sigma^0/\sqrt{2} & \Sigma^+ \\ \Sigma^- & -\Sigma^0/\sqrt{2} \end{pmatrix}$ .  
 $q\bar{q} \rightarrow Z^*/\gamma^* \rightarrow \Sigma^+ \Sigma^-$   
 $q\bar{q}' \rightarrow W^* \rightarrow \Sigma^+ \Sigma^0$

(Bajc, Senjanovic, Nemevsek)



# Accessing seesaw scale in low energy expts

- Possible new operators from new physics:

Type I →

$$\delta\mathcal{L}^{d=6} = c_{\alpha\beta}^{d=6} \left( \overline{\ell_{L\alpha}} \tilde{\phi} \right) i\not{\partial} \left( \tilde{\phi}^\dagger \ell_{L\beta} \right) \quad \mathcal{E}^2$$

→ Leads to non-unitarity of PMNS: SBL tests possible for  $\epsilon > 10^{-2}$

Current bounds:

$$|\eta| < \begin{pmatrix} 2.0 \times 10^{-3} & 3.5 \times 10^{-5} & 8.0 \times 10^{-3} \\ 3.5 \times 10^{-5} & 8.0 \times 10^{-4} & 5.1 \times 10^{-3} \\ 8.0 \times 10^{-3} & 5.1 \times 10^{-3} & 2.7 \times 10^{-3} \end{pmatrix}$$

- Type II: 
$$\delta\mathcal{L}_{4F} = \frac{1}{M_\Delta^2} \left( \overline{\tilde{\ell}_L} Y_\Delta \vec{\tau} \ell_L \right) \left( \overline{\tilde{\ell}_L} \vec{\tau} Y_\Delta^\dagger \tilde{\ell}_L \right)$$

- Type III: 
$$\delta\mathcal{L}^{d=6} = c_{\alpha\beta}^{d=6} \left( \overline{\ell_{L\alpha}} \vec{\tau} \tilde{\phi} \right) i\not{\partial} \left( \tilde{\phi}^\dagger \vec{\tau} \ell_{L\beta} \right), \quad \sim \mathcal{E}^2$$

# Seesaw and new gauge forces:

- Seesaw strongly suggests gauge forces
- Type I : why seesaw scale so far below Planck scale:
- Inverse seesaw case:

- **Why** 
$$\begin{pmatrix} 0 & hv_{wk} & 0 \\ hv_{wk} & 0 & M \\ 0 & M & \mu \end{pmatrix}$$
 **why not** 
$$\begin{pmatrix} 0 & hv_{wk} & h'v_{wk} \\ hv_{wk} & M' & M \\ h'v_{wk} & M & \mu \end{pmatrix}$$

- **Local B-L** symmetry --a plausible answer !
- Collider profile of seesaw undergoes drastic change !!

# What Gauge Symmetry ?

- Standard model: gauge sym.
- Fermions:

$$SU(2)_L \times U(1)_Y$$

$$m_\nu = 0$$



$$\begin{pmatrix} u_L \\ d_L \end{pmatrix} \quad \begin{matrix} u_R \\ d_R \end{matrix}$$

$$\begin{pmatrix} \nu_L \\ e_L \end{pmatrix} \quad \begin{matrix} \phantom{\nu_R} \\ e_R \end{matrix}$$

- $N_R \rightarrow$  Gauge group:**

$$SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$$

$$\begin{pmatrix} u_L \\ d_L \end{pmatrix} \quad \overset{P}{\Leftrightarrow} \quad \begin{pmatrix} u_R \\ d_R \end{pmatrix}$$

$$\begin{pmatrix} \nu_L \\ e_L \end{pmatrix} \quad \overset{P}{\Leftrightarrow} \quad \begin{pmatrix} \nu_R \\ e_R \end{pmatrix}$$

- New**



$$W_L^\pm$$

$$W_R^\pm$$

$$Z, Z', \gamma$$

# Bounds on LR Scale

- Most stringent bounds come from CP viol. Observables  $\varepsilon, \varepsilon', d_n^e$  ; depends on how CP is introduced:

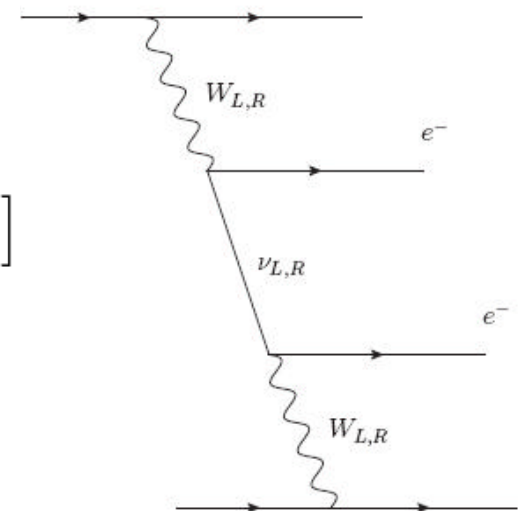
$$M_{W_R} \geq 2.5 \text{TeV} \text{ to } 4 \text{TeV} \quad \text{Zhang et al; Maiezza et al.}$$

- New contributions to nu-less double beta decay  
**(Type I)** (RNM, 86; Hirsch, Klapdor, Panella 96)

$$\rightarrow m_{W_R} \geq 1.1 \left( \frac{\langle m_N^{(V)} \rangle}{1 \text{TeV}} \right)^{(-1/4)} [\text{TeV}]$$

**From Ge76:**

- Collider limits (D0,CDF)  $> 750 \text{ GeV}$



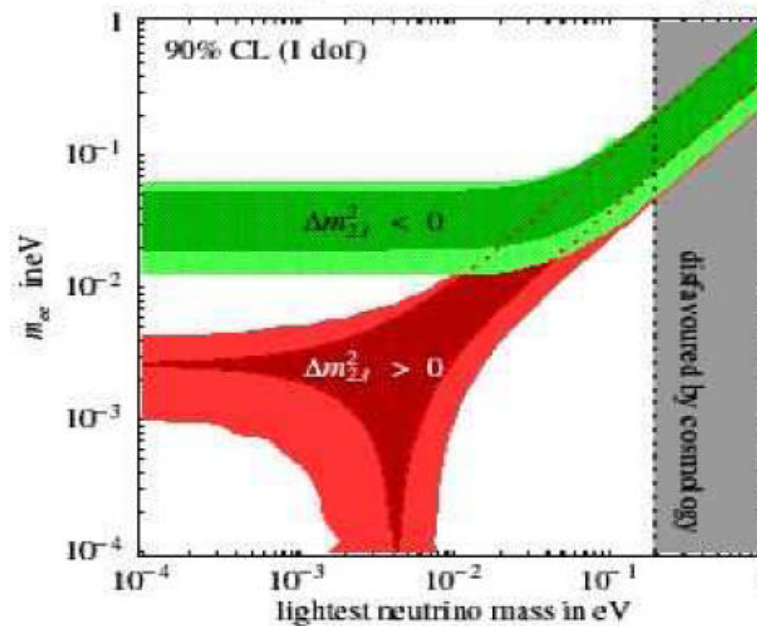
# TeV WR signal from $\beta\beta_{0\nu}$

## ■ Nu contribution:

Inverse hierarchy

Normal hierarchy

(Ferruglio, Strumia, Vissani;  
Petcov, Passcoli, Bilenky)



## Punch line:

■ Suppose long baseline  $\rightarrow \Delta m_{31}^2 > 0$

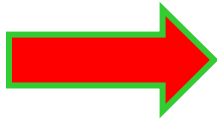
■ and nonzero signal for  $\beta\beta_{0\nu}$

$\rightarrow$  could be a signal of TeV WR and type I

# Theory insights into seesaw scale

- **Upper limits: Supersymmetry**

Supersymmetry restricts potential for SUSY LR type I seesaw



$$M_{W_R} < \frac{M_{susy}}{f}$$

Kuchimanchi, RNM, 95

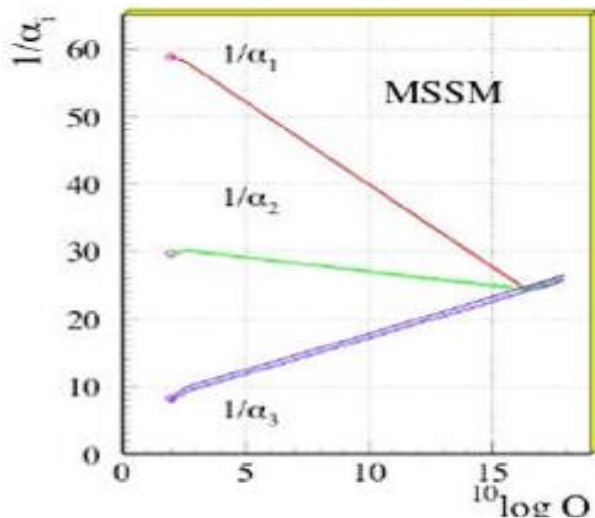
- **Sneutrino or KK Nu<sub>R</sub> dark matter →**

$$M_{W_R} < \text{few TeV}$$

Kong, Matchev, Nasri;; Hsieh, Nasri, RNM

# Is Grand unification idea any help ?

- Gauge couplings unify with TeV scale SUSY-MSSM



→ Suggests grand unified Th.

- Connects quark and lepton physics-
- SU(5) does not help- Seesaw scale free parameter: (proton decay ?)



# SO(10) just right:

- {16}- spinor for all matter  
-includes RH neutrino

$$\begin{pmatrix} u & u & u & \nu \\ d & d & d & e \end{pmatrix}_{L,R}$$

- Q-I unification predicts

$$m_D \approx m_t$$

- Type I case: seesaw scale

$$M_R \approx 10^{14} \text{ GeV}$$

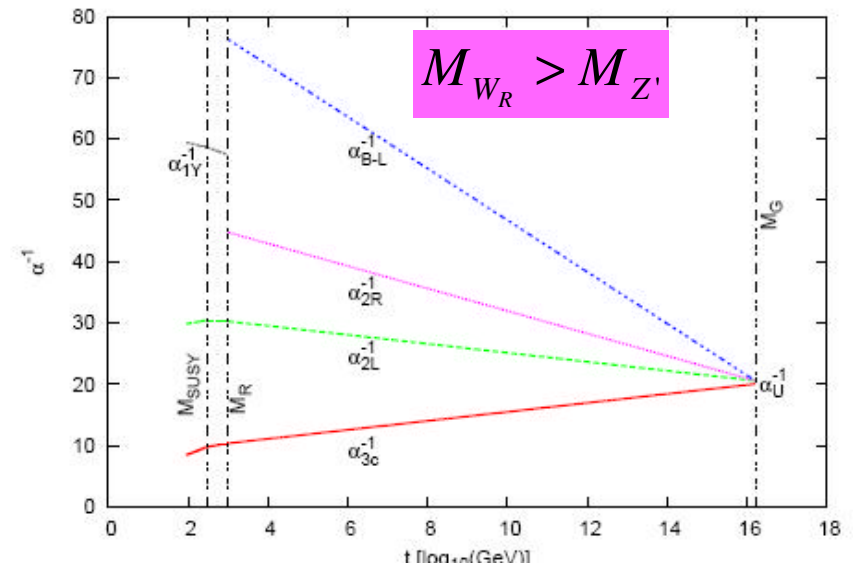
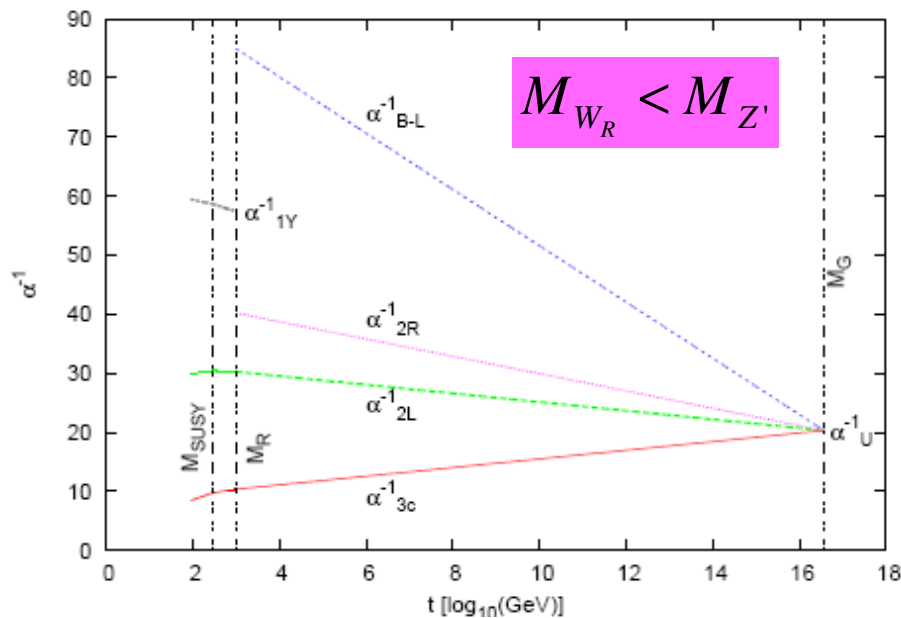
- **Very appealing picture:** predictive with extra global family symmetries;
- **TeV scale for type I does not unify:**

(Parida, Raichoudhuri, Majee, Sarkar; Kopp, Lindner, Niro, Underwood)



# Inverse seesaw in SO(10) - LHC accessible

- New result!** Inverse seesaw does unify –TeV  $W_R Z'$



- $$SO(10) \xrightarrow{M_G} 3_c 2_L 2_R 1_{B-L} \xrightarrow{M_R} 3_c 2_L 1_Y (\text{MSSM}) \xrightarrow{M_{\text{SUSY}}} 3_c 2_L 1_Y (\text{SM}) \xrightarrow{M_Z} 3_c 1_Q$$

$$M_U \cong 10^{16} \text{ GeV}; M_{BL,R} \cong \text{TeV}$$

(Dev, RNM, 09; PRD; arXiv 1003:6102);

# Leptogenesis and TeV WR

- Formula for baryon asym

$$\eta_B \simeq 10^{-2} \sum_{i,\alpha} \epsilon_{i\alpha} \kappa_{i\alpha}$$

- Resonant lepto:  $\epsilon \sim \frac{h_D^2 \delta}{\Delta M / M}$  ;  $\kappa \sim \frac{D}{D+S} \exp[-W_{L=2}^{ID,\dots}]$
- Type I case:**  $h_D^2 \sim 10^{-12}$ ;  $\rightarrow \frac{\Delta M}{M} \sim 10^{-12}$ ; S can be  $\gg D$
- Washout depends on WR mass: **MWR > 17 TeV** (Frere, Hambye, Vertoggen)
- Inverse seesaw**  $h_D^2 \sim 10^{-2} \rightarrow D \gg S$ ;  $W_{L=2} \sim \frac{\mu}{M} \sim 10^{-7} \rightarrow \kappa \sim 1$
- Not so sensitive dependence on WR. **MWR > 1.5 TeV** (Blanchet, Dev, RNM, to appear)

# Type I vs Inverse left-right seesaw at LHC

- WR and  $Z'$ :  $u\bar{d} \rightarrow W_R \rightarrow l^+ N$      $u\bar{u} \rightarrow Z' \rightarrow NN$   
(Keung, Senjanovic; Han, Perez, Huang, Li, Wang; Del Aguila, Aguilar-Saavedra; de Blas, Azuelos,
  - N-decay: (a)  $\nu N$  mixing  $\propto \mathcal{E}$  / or (b) exchange  $W_R$
  - **type I** :  $\mathcal{E} \ll 10^{-3}, M_{W_R} < 4\text{TeV}$  negligible;     $N \rightarrow l^\pm jj$
  - **Signal**  $pp \rightarrow l^\pm l^\pm jj + X$
  - **Inv. Seesaw**  $\mathcal{E} > 10^{-3}$      $N \rightarrow l^- jj, l^- l^+ \nu$      $pp \rightarrow lll\nu + X$
- LHC reach < 4 TeV**
- Determination of chirality: **tb decay mode** (Gopalakrishna, Han, Lewis, Si, Zhou)

## 2. Understanding Large mixings

### Issue 1. Large lepton mixings: why ?

#### Leptons:

$$U_{\text{PMNS}} \approx \begin{pmatrix} \frac{\sqrt{6}}{3} & \frac{\sqrt{3}}{3} & 0 \\ -\frac{\sqrt{6}}{6} & \frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \\ \frac{\sqrt{6}}{6} & -\frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \end{pmatrix} + \delta U$$

### Issue 2. How to have a unified understanding of quark and lepton mixings

$$U_{\text{CKM}} = \begin{pmatrix} 1 & \lambda & \lambda^3 \\ -\lambda & 1 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix}$$



# Lepton mixings

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## **(i) Family symmetries** :

$$S_{2(\mu-\tau)} \subset S_3, S_4, A_4, \Delta(3n^2), \dots$$

- **Generically leads to TBM + small corrections**

(Talks by Morisi and Hagedorn)

## **(ii) Dynamical e.g. RH neutrino dominance** (King)

generically predicts large  $\theta_{13}$  ;

→ no explanation of TBM;

# A strategy for flavor unification

**A**

$$M_u = M_0 + \delta_u$$

$$M_d = rM_0 + \delta_d$$

$$M_l = rM_0 + \delta_l$$

(Dutta, Mimura, RNM'PRD-09)

$$m_\nu = f\nu_L$$

$$\delta_{u,d,l} \ll M_0$$

→ **Anarchic**  $M_0, \mathbf{f}$  → **quark mixings small while lepton mixings large** +  $m_b \cong m_\tau$

**Natural** in SO(10)+ type II seesaw

**B**

**Rank 1  $M_0$**  → **explains mass hierarchies**

# An $S_4 \times SO(10)$ - example

- Symmetries make theory predictive: e.g.  $S_4$

- Solar mass

$$\frac{m_{solar}}{m_{atm}} \cong \lambda \cong \theta_c$$

Dutta, Mimura, RNM arXiv:0911.2242 , JHEP

- Bottom-tau:

$$m_b \approx m_\tau$$

and

$$m_\mu = -3m_s$$

- Leading order PMNS- Tri-bi-maximal

Corrections: Testable

$$\theta_{13} = \frac{\theta_c}{3\sqrt{2}} \cong 0.05$$

Bjorken, King, Pakvasa Ferrandis; Chen, Mahanthappa

- Prediction for atmospheric angle:  $\theta_{23} \cong 35^\circ - 40^\circ$

- Current fit:  $\theta_{23}^0 = 42.8_{-2.9-7.3}^{+4.7+10.7}$

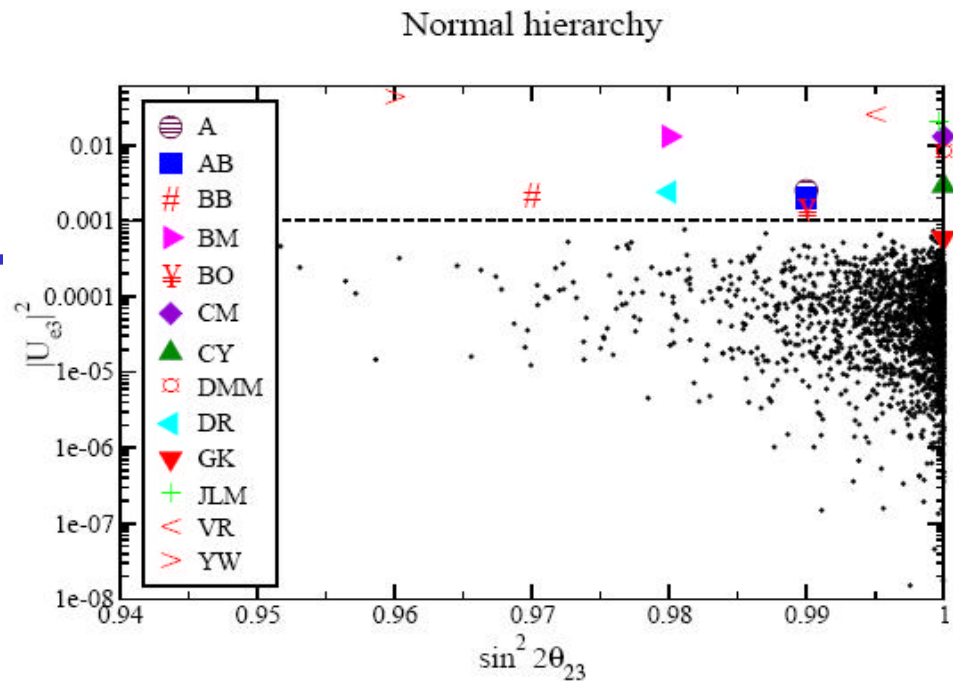
(Gonzalez-Garcia, Maltoni, Salvado,2010)

# Generic GUT predictions

- Normal hierarchy  $m_1 < m_{solar} < m_{atm}$

- $\theta_{13}$  predictions for GUT vs non-GUT

(Compilation by Albright, Rodejohann)







## 3. Schizophrenic neutrinos

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- Common assumption: Neutrinos are either Dirac or Majorana
- “Neutrino---neu-ter---(not either)”; E. Lisi, Neutrino 2010 summary talk
- **Could it be that neutrinos are neither fully Dirac or fully Majorana but little bit of both i.e. Schizophrenic ?**

# How could that be ?

- Experiments →  $U_{PMNS}$ , which relates mass to flavor eigenstates

$$V_{flavor} = U_{PMNS} V_{mass}$$

- May be some **mass eigenstates** are Dirac –others Majorana; **each flavor will then be both Dirac and Majorana-** mixings and masses are unaffected.

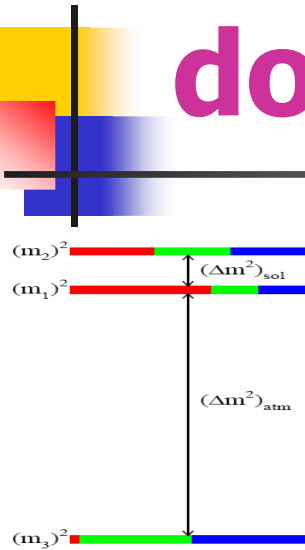
- Specific example:  $v_2 \equiv \frac{1}{\sqrt{3}}(v_e + v_\mu + v_\tau)$  Dirac -others Majorana

- Why this combo?  $S_3$  sym permuting:  $(L_e, L_\mu, L_\tau)$

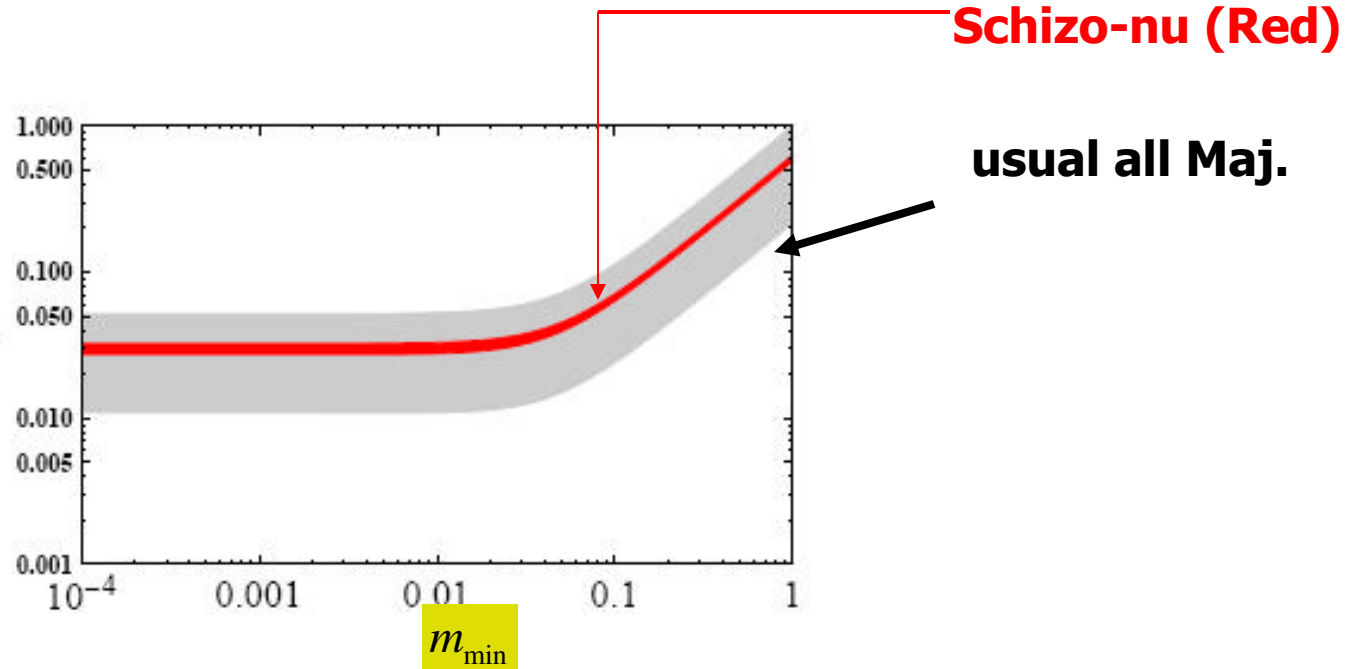
$$\{\mathbf{1}\}: \frac{1}{\sqrt{3}}(L_e + L_\mu + L_\tau) \text{ Dirac; } \{\mathbf{2}\}: \left[ \frac{1}{\sqrt{2}}(L_\mu - L_\tau), \frac{1}{\sqrt{6}}(2L_e - L_\mu - L_\tau) \right] \text{ Maj}$$

(Allahverdi, Dutta, RNM, arXiv:1008.1232)

# Implications for Nu-less double beta decay-**IH**



$m_{ee}$



- $M_{\min}^{\beta\beta} = 17 \text{ meV}$  (usual IH) vs  $34 \text{ meV}$  (Schizo-nu IH) –  
→ schizo-nu easier to rule out if LBL  $\rightarrow$  IH

# Why ? Inflation → Schizophrenia

- **Susy B-L theory for seesaw → 3 RH neutrinos**

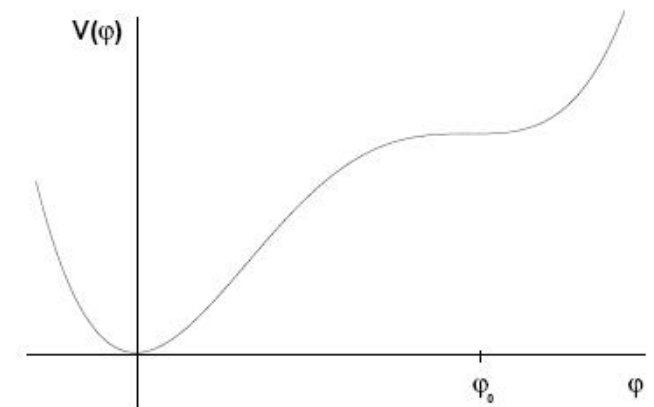
→ To fit neutrino observations, we only need two RH nu's- **so what is the third one (N3) doing ?**

→ **Could its superpartner be the inflaton ?**

→ **Inflation requires  $\sim$ flat potential,**

→ **In the B-L theory ( $\tilde{N} + \tilde{L} + H_u$ )  
is one provided  $h_D$  is tiny.**

(Allahverdi, Kusenko, Mazumdar;  
Allahverdi, Dutta, Mazumdar)





# Implications for neutrinos

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- **How small is  $h_D$  ?**
- **COBE**  $\frac{\delta\rho}{\rho} \sim 10^{-5} \rightarrow h_D \sim 10^{-12}$  **for only one RH neutrino; just right for a Dirac nu:**
- **That RH neutrino cannot mix with others to maintain flatness.**
- Immediately suggests this scenario for nu masses



# Conclusion:

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- **Seesaw: Compelling big picture idea for neutrino masses;**

**(i) LHC can discover TeV scale seesaw;  
even in SO(10) models; Must it discover ? NO**

**(ii) Most appealing neutrino models are based  
on type II+ SO(10) GUT seesaw; No LHC signal;  
Can also unify flavor with extra sym.**

**(iii) An inflation inspired schizophrenic  
neutrino scenario.**

# Other advantages of a Low seesaw scale

- **Leptogenesis**- : High scale leptogenesis + hierarchical NR (as in type I or II SO(10) →

$$M_N \geq 10^9 \text{ GeV}$$

vs

$$T_R < 10^6 - 10^9 \text{ GeV}$$

adequate leptogenesis

gravitino reheat constraint

→ Low scale resonant leptogen as alternative:

- **Suppression of proton decay**: High scale seesaw SO(10) with 16-Higgs has proton decay problem from operator;

→ suppressed for TeV B-L breaking.

$$\frac{(16_m)^3 16_H}{M_{Pl}}$$

# SO(10) with type II seesaw → realization of flavor ansatz

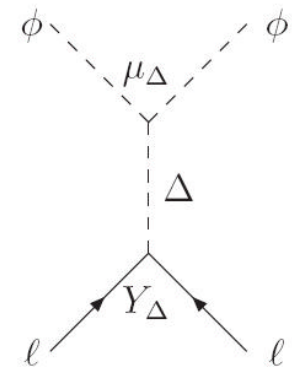
- SO(10) → fermion mass formulae:

$$\begin{aligned} Y_u &= h + r_2 f + r_3 h' \\ Y_d &= r_1(h + f + h') \\ Y_e &= r_1(h - 3f + c_e h') \end{aligned}$$

(Babu, Mohapatra'92)

$$m_\nu \cong f \nu_\Delta$$

Bajc, Senjanovic, Vissani'03



- For  $f, h' \ll h$ , → yields ansatz part A at **Mu**;
- Rank (part B) from flavor symmetry:
- Rank solves the proton decay problem also



# LHC accessible non-seesaw models

- Seesaw scenarios based on operator:  $\frac{LHLH}{M}$
  - Higher the dimension of operator, lower the scale:  
e.g.  $LLHH(H^\dagger H)/M^3$  (Babu, Nandi and Tavartkiladze)
- TeV scale theory; predicts a field  $\Phi^{+++}$   $M < \text{TeV}$   
observable via decays:  $\Phi^{+++} \rightarrow W^+W^+W^+, W^+e^+e^+$

# Inverse Seesaw: details

- Rate plots:

