

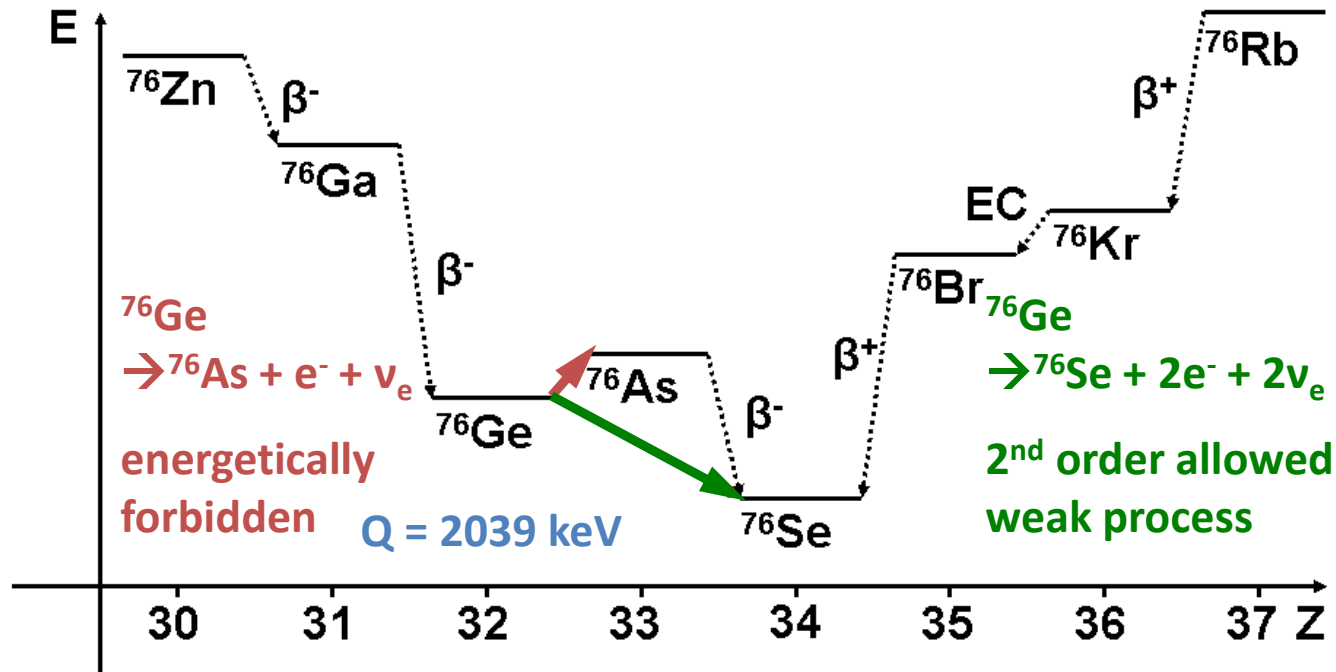


# Neutrinoless double beta decay: Experimental challenges

Konstantin Gusev  
JINR, Dubna

NOW 2018  
September 2018 – Rosa Marina (Ostuni, Italy)

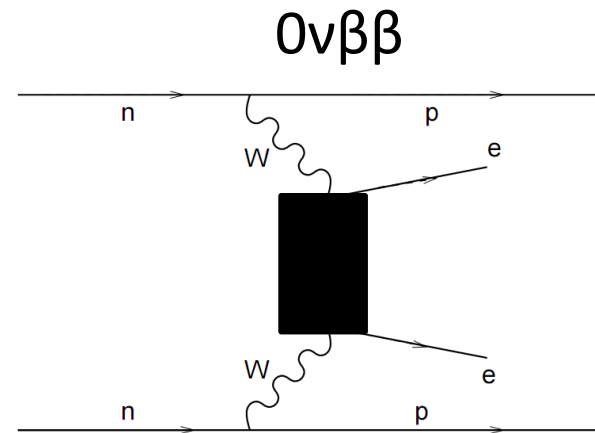
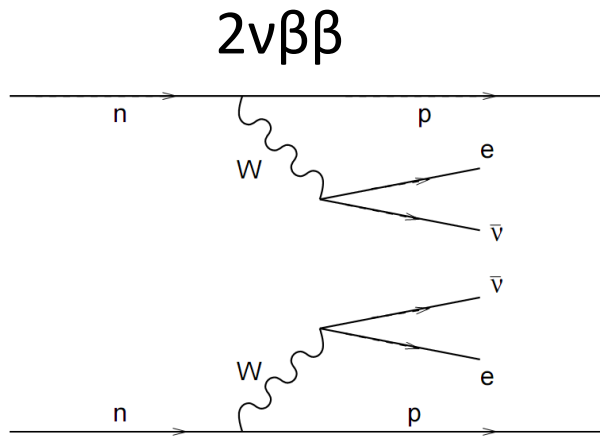
# Short intro in $0\nu\beta\beta$ -decay



- ✓ In 1936 Maria Göppert-Mayer noted, that in some even-even nuclei the single  $\beta$ -decay is energetically forbidden whereas the simultaneous but independent  $\beta$ -decay of two nucleons (so-called double beta decay) is allowed

# Short intro

## Why $(0\nu)\beta\beta$ -decay?



- violates lepton number? **NO**
- forbidden in SM? **NO**

- violates lepton number? **YES!**
- forbidden in SM? **YES!**

- but half life is  $10^{10}$  longer than the age of the universe, however already observed!

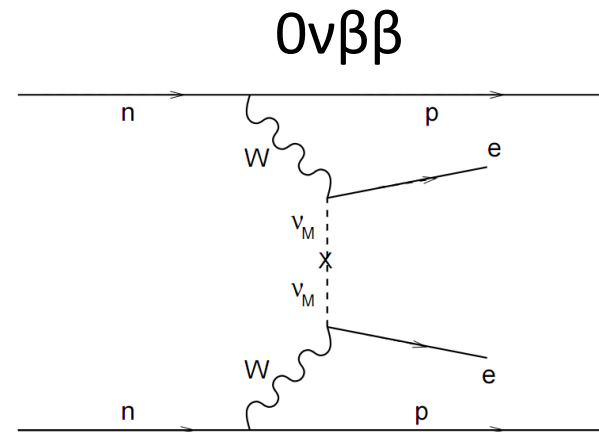
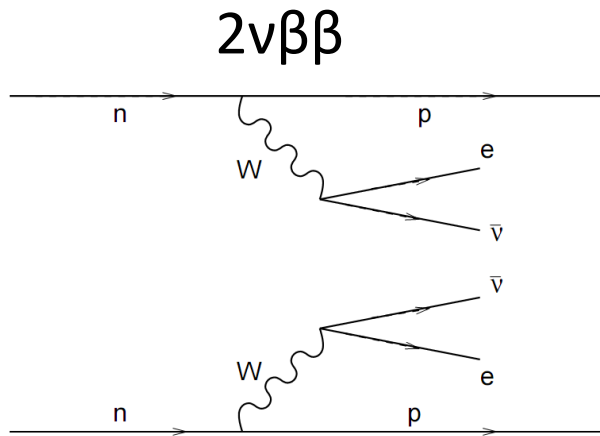
$^{76}\text{Ge}: T_{1/2}^{2\nu} = 1.92 \times 10^{21}\text{yr}$

**New Physics!**

- $\nu$  has **Majorana** mass component

# Short intro

## Why $(0\nu)\beta\beta$ -decay?



- violates lepton number? **NO**
- forbidden in SM? **NO**

- violates lepton number? **YES!**
- forbidden in SM? **YES!**

- but half life is  $10^{10}$  longer than the age of the universe, however already observed!

- $\nu$  has **Majorana** mass component
- **IF** light neutrino exchange

$^{76}\text{Ge}: T_{1/2}^{2\nu} = 1.92 \times 10^{21}\text{yr}$

**Access to  $\nu$  mass scale**

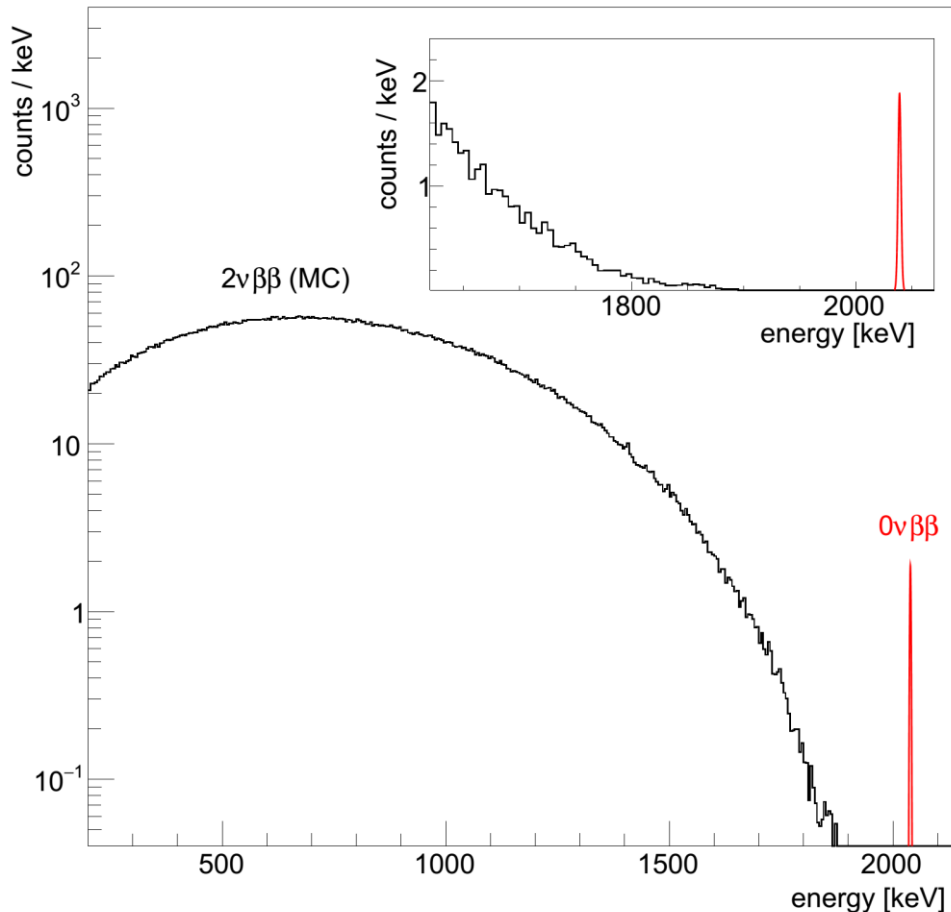
**New Physics!**



# Short intro

# What are we measuring?

Summed electron spectrum ( $^{76}\text{Ge}$ ):



**0νββ:**

Sharp peak at Q-value of the decay

$$T_{1/2}^{0\nu} > 10^{25} \text{yr}$$

**2νββ:**

Continuous spectrum

$$T_{1/2}^{2\nu} \sim 10^{21} \text{yr}$$



Background for 0νββ

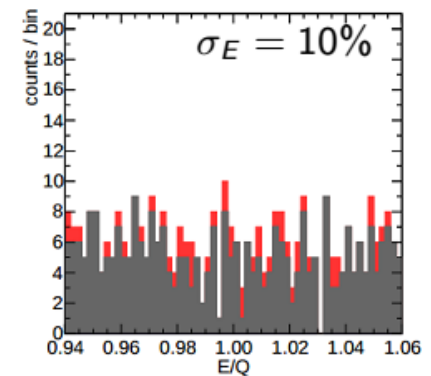
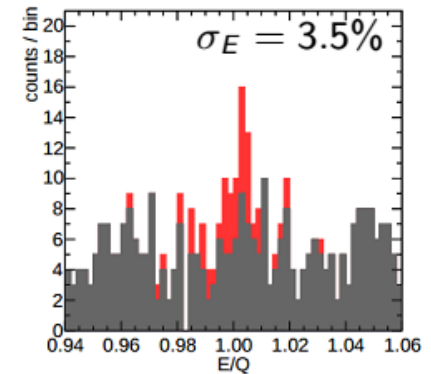
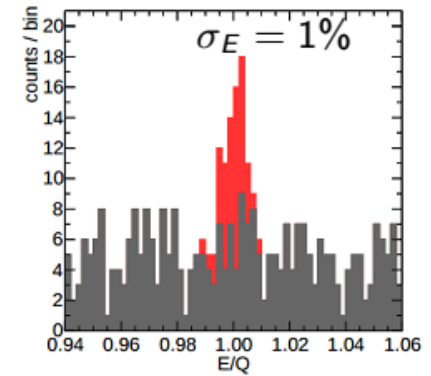
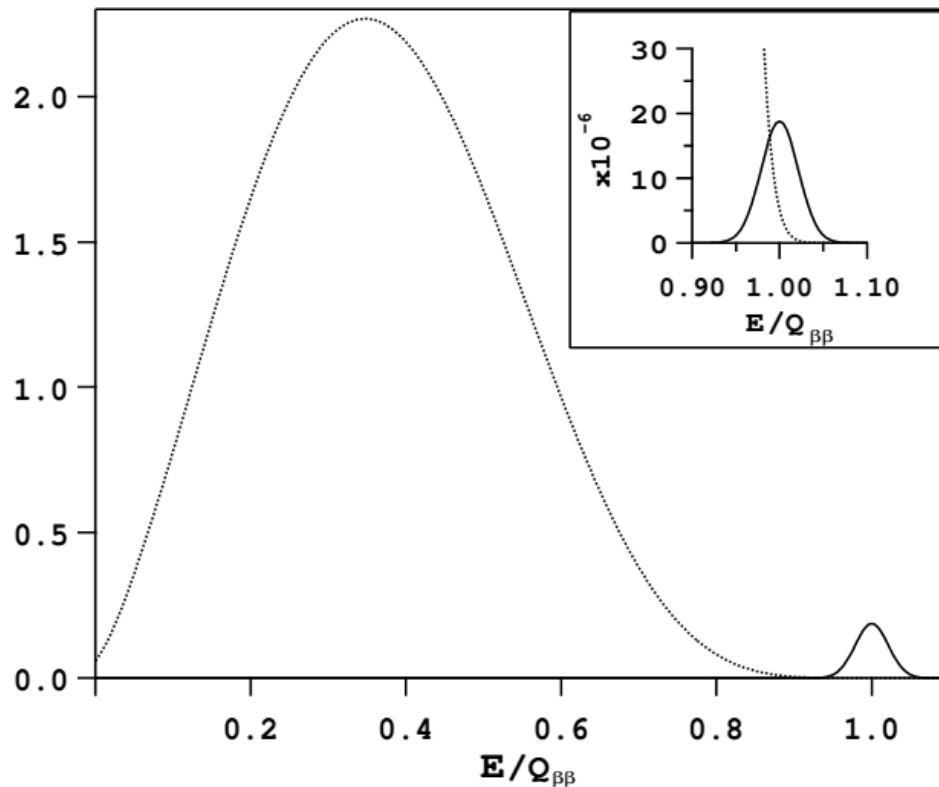


Energy resolution essential

# Short intro

# What are we measuring?

✓ Resolution remains essential due to  $2\nu\beta\beta$



# Short intro

## How to measure?

### Experimental sensitivity:

- **Zero** background:

$$T_{1/2}^{0\nu} \propto M t$$

- Non-zero background:

$$T_{1/2}^{0\nu} \propto \sqrt{\frac{M t}{\Delta E BI}}$$

$M t$  – exposure (kg yr)

$\Delta E$  – energy resolution (keV)

$BI$  – background index (counts/keV kg yr)

Isotope	$G^{0\nu}$ ( $10^{-14}$ yr)	$Q$ (keV)	Nat. ab. (%)
$^{48}\text{Ca}$	6.3	4273.7	0.187
$^{76}\text{Ge}$	0.63	2039.1	7.8
$^{82}\text{Se}$	2.7	2995.5	9.2
$^{100}\text{Mo}$	4.4	3035.0	9.6
$^{130}\text{Te}$	4.1	2530.3	34.5
$^{136}\text{Xe}$	4.3	2461.9	8.9
$^{150}\text{Nd}$	19.2	3367.3	5.6

enrichment required except for  $^{130}\text{Te}$ ,  
not (yet) possible for all, costs differ

- ✓ Target mass and detector efficiency as high as possible
- ✓ **“Zero-background”** to have linear increase of sensitivity vs exposure

# Short intro

## What about mass?

Effective Majorana neutrino mass contributes in the decay rate:

Phase space factor

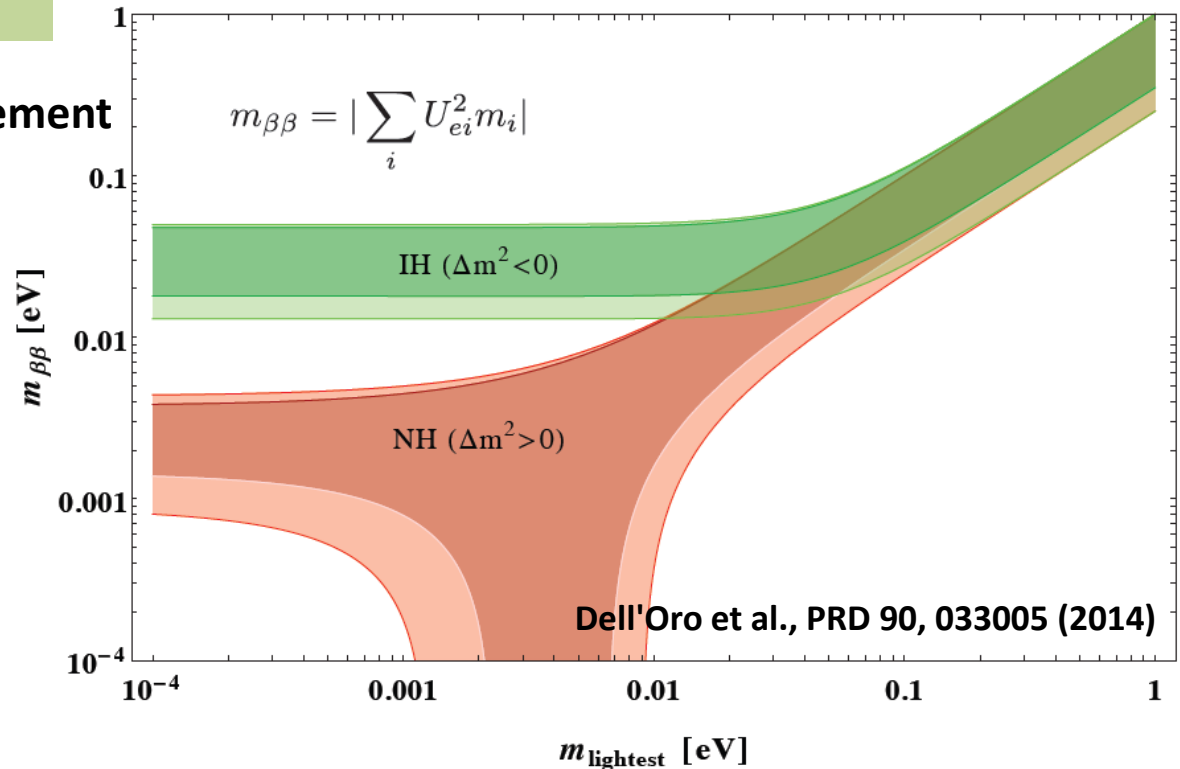
Effective Majorana neutrino mass

$$\frac{1}{T_{1/2}^{0\nu}} = G^{0\nu} |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$

$0\nu\beta\beta$  decay rate

Matrix element  
(NME)

$$\langle m_{\beta\beta} \rangle = \left| \sum_i U_{ei}^2 m_i \right|$$

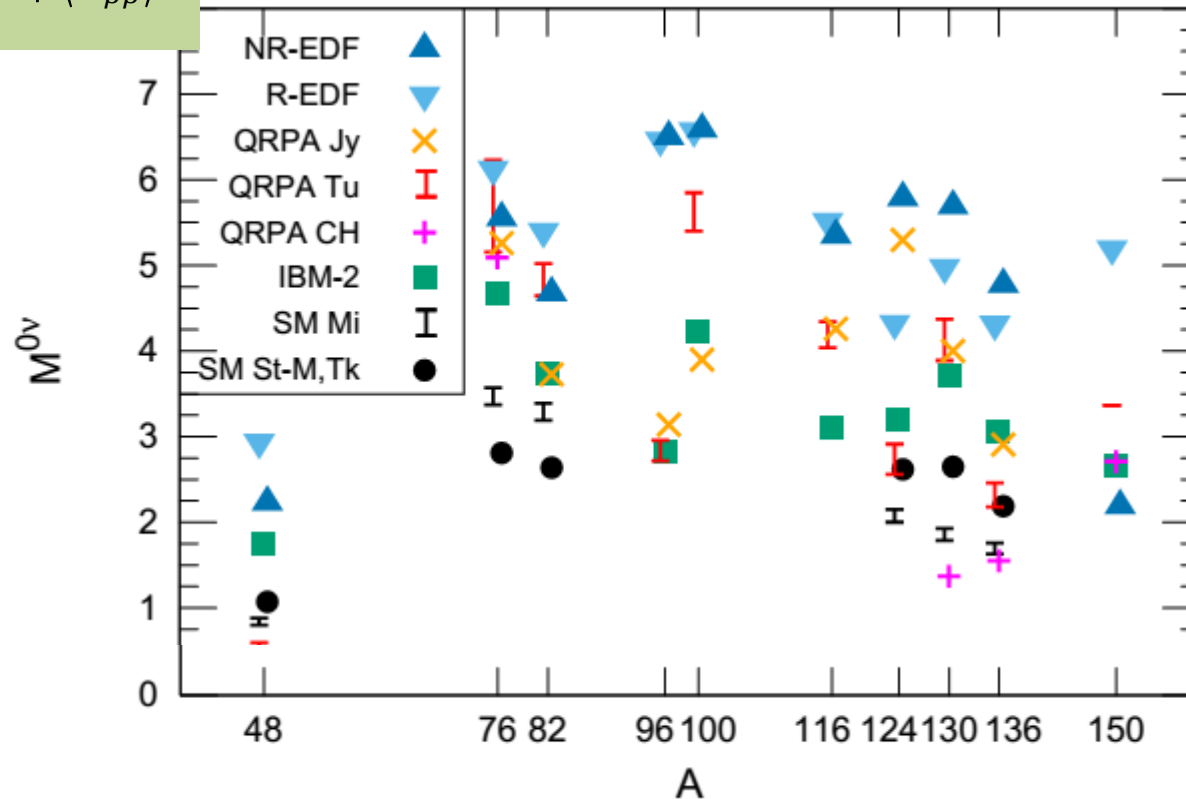




# Short intro

## NME

$$\frac{1}{T_{1/2}^{0\nu}} = G^{0\nu} |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$



Engel & Menéndez  
arXiv: 1610.06548v2

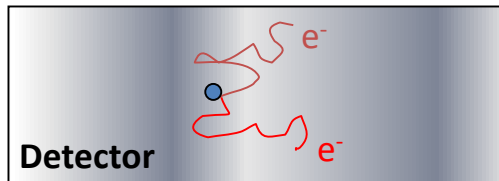
✓ No preferred isotope from Nuclear Physics ( $G^*M$ )

See today's talk  
by Fedor Simkovic

# Short intro

## Two experimental approaches

### Source = Detector



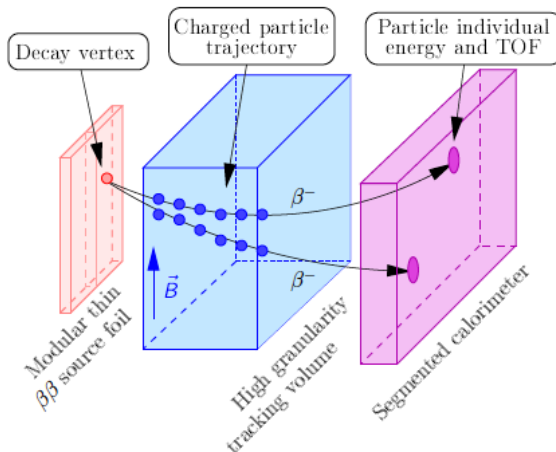
GERDA, MJD, CUORE,  
EXO, Kamland-Zen,  
SNO+, ...

- + High detection efficiency
- + Large target mass possible

± Reconstruction of event topologies

– Restricted number of isotopes

### Source ≠ Detector



SuperNEMO

- + Reconstruction of event topologies
- + Coincidence scheme
  - zero background
- + No restriction on isotopes

– Difficult to obtain large masses

# $0\nu\beta\beta$ -experiments

## Now

Gas or liquid <i>Easy to get huge mass</i>	TPCs	EXO-200 NEXT-10
	Liquid scintillators	KamLAND-Zen
Crystal <i>Energy resolution</i>	Bolometers	CUORE CUPID-0,-Mo AMoRE
	Ge-detectors	GERDA MJD
Source $\neq$ Detector		

*Adapted from A. Giuliani, Neutrino2018*

# $0\nu\beta\beta$ -experiments

## Soon

**Now**

Gas or liquid <i>Easy to get huge mass</i>	TPCs	EXO-200 NEXT-10	NEXT-100 PANDA-X-III
	Liquid scintillators	KamLAND-Zen	KZ-800 SNO+ phase I
Crystal <i>Energy resolution</i>	Bolometers	CUORE CUPID-0,-Mo AMoRE	AMoRe II
	Ge-detectors	GERDA MJD	LEGEND-200
Source $\neq$ Detector			SuperNEMO

*Adapted from A. Giuliani, Neutrino2018*

# $0\nu\beta\beta$ -experiments

## Future

		Now	Soon	
Gas or liquid <i>Easy to get huge mass</i>	TPCs	EXO-200 NEXT-10	NEXT-100 PANDA-X-III	nEXO NEXT-2.0 PANDAX-III 1t
	Liquid scintillators	KamLAND-Zen	KZ-800 SNO+ phase I	KamLAND2-Zen SNO+ phase II
Crystal <i>Energy resolution</i>	Bolometers	CUORE CUPID-0,-Mo AMoRE	AMoRe II	CUPID
	Ge-detectors	GERDA MJD	LEGEND-200	LEGEND-1000
Source $\neq$ Detector			SuperNEMO	

*Adapted from A. Giuliani, Neutrino2018*

# $0\nu\beta\beta$ -experiments

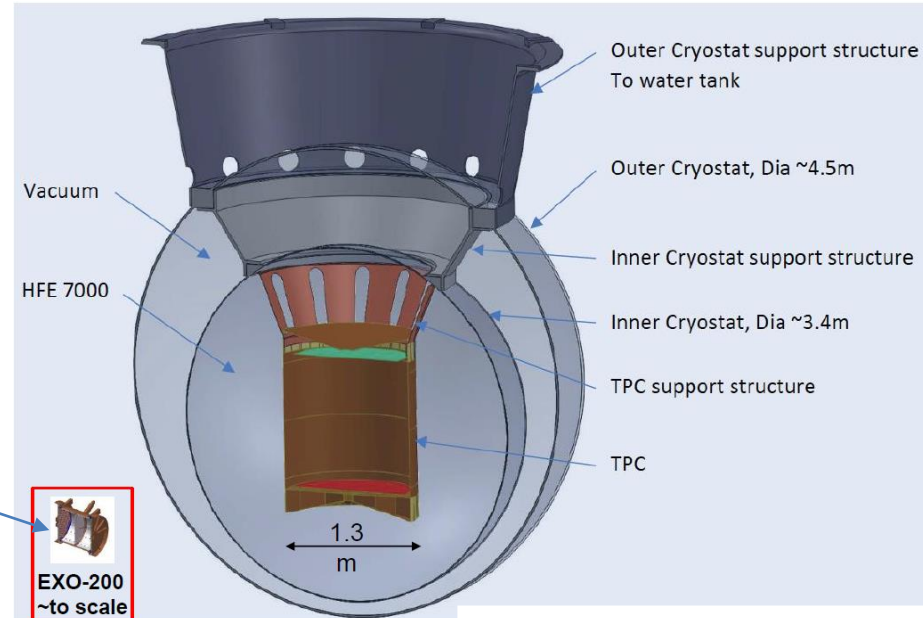
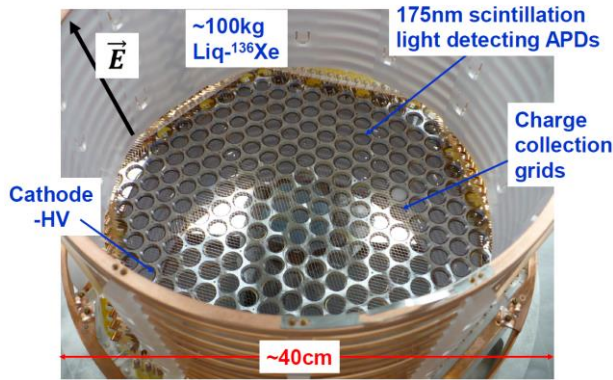
## Now 2018

		Now	Soon	Future
Gas or liquid <i>Easy to get huge mass</i>	TPCs			See today's talk by <b>Yoshihito Gando</b>
	Liquid scintillators	<b>KamLAND-Zen</b>	<b>KZ-800</b> <b>SNO+ phase I</b> See talk on Sept 11 by <b>Edward Leming</b>	<b>KamLAND2-Zen</b> <b>SNO+ phase II</b>
Crystal <i>Energy resolution</i>	Bolometers	<b>CUORE</b> <b>CUPID-0,-Mo</b>	See today's talk by <b>Paolo Gorla</b> <b>CUPID</b> See today's talk by <b>Nicola Casali</b>	
	Ge-detectors	<b>GERDA</b>	See today's talk by <b>Christoph Wiesinger</b>	
Source $\neq$ Detector				

# $0\nu\beta\beta$ -decay with TPCs EXO-200 and nEXO



The EXO-200 liquid  $^{136}\text{Xe}$  Time Projection Chamber

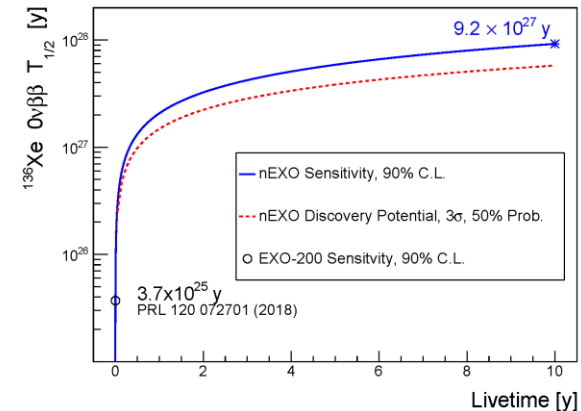


Method	Liquid TPC
Location	WIPP, USA
Isotope	$^{136}\text{Xe}$
$T_{1/2}$ sensitivity	$> 3.7 \cdot 10^{25}$ yr (90% CL)
Limit	$> 1.8 \cdot 10^{25}$ yr (90% CL)

PRL 120 072701 (2018)

From EXO-200 to nEXO:

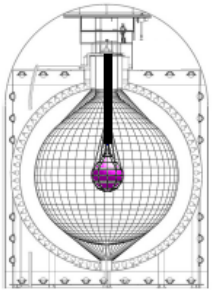
- ✓ 150 kg  $\rightarrow$  5000 kg
- ✓ WIPP  $\rightarrow$  SNOLAB
- ✓ 3.7  $\rightarrow$  920 ( $\times 10^{25}$ ) yr



EXO-200 will finish data taking in 2018

# $0\nu\beta\beta$ -decay with liquid scintillators

## KamLAND-Zen, 800, 2-Zen



### Past

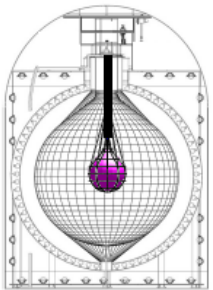
#### KamLAND-Zen 400

320-380 kg of Xenon  
Data taking 2011 ~ 2015



Method	Xe-loaded LS
Location	Kamioka, JAPAN
Isotope	$^{136}\text{Xe}$
$T_{1/2}$ sensitivity	$> 5.6 \cdot 10^{25}$ yr (90% CL)
Limit	$> 1.1 \cdot 10^{26}$ yr (90% CL)

*PRL 117 109903 (2016)*



### Present

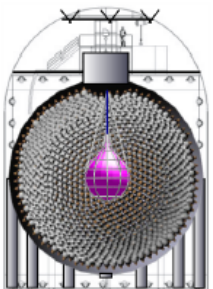
#### KamLAND-Zen 800

~750 kg of Xenon  
DAQ to start in this year



Status 2018:

- ✓ New ballon installed
- ✓ Filled in May 2018 with dummy LS
- ✓ Will be replaced with Xe-loaded LS



### Future

#### KamLAND2-Zen

~1 ton of  $^{136}\text{Xe}$   
Better energy resolution



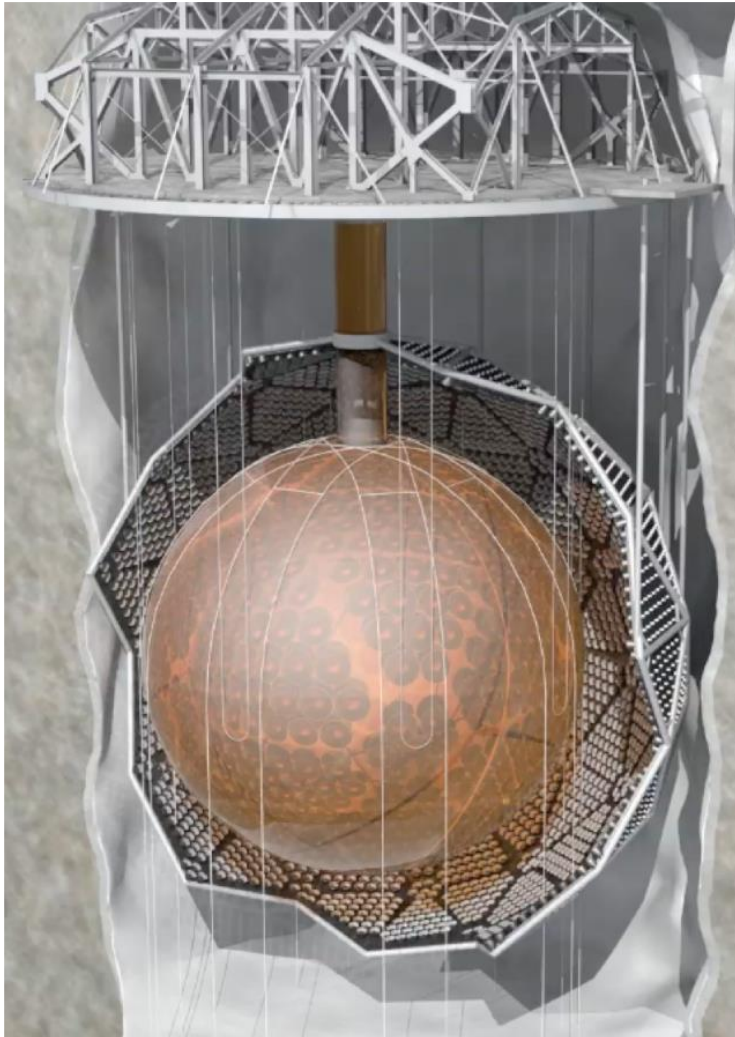
- ✓ Brighter LS
- ✓ New PMTs
- Better energy resolution

See today's talk  
by Yoshihito Gando



# $0\nu\beta\beta$ -decay with liquid scintillators

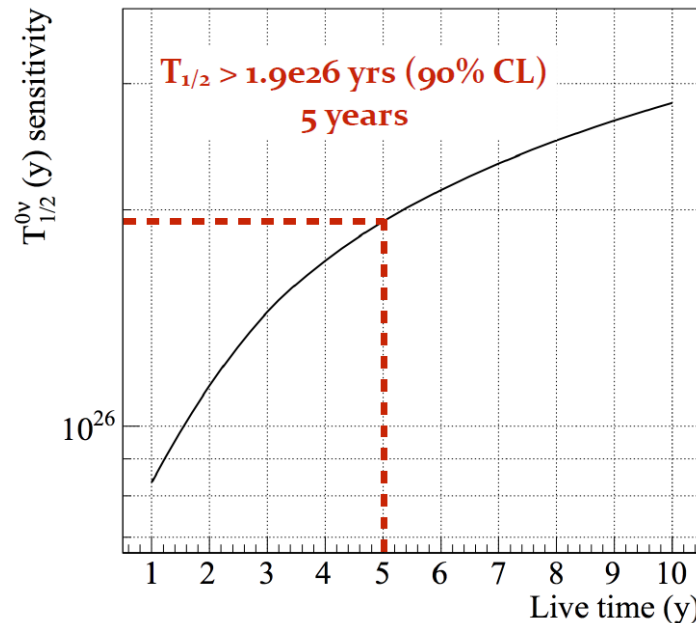
## SNO+



- ✓ SNOLAB, Ontario
- ✓ 780 ton LAB/PPO (2g/L) in 6m radius acrylic vessel
- ✓ ~9400 PMTs at 8.5m

Phased implementation:

- ✓ Water phase → ongoing
- ✓ Pure scintillator phase → LS fill in July 2018
- ✓ Loaded scintillator phase → 0.5%  $^{130}\text{Te}$  loading in 2019

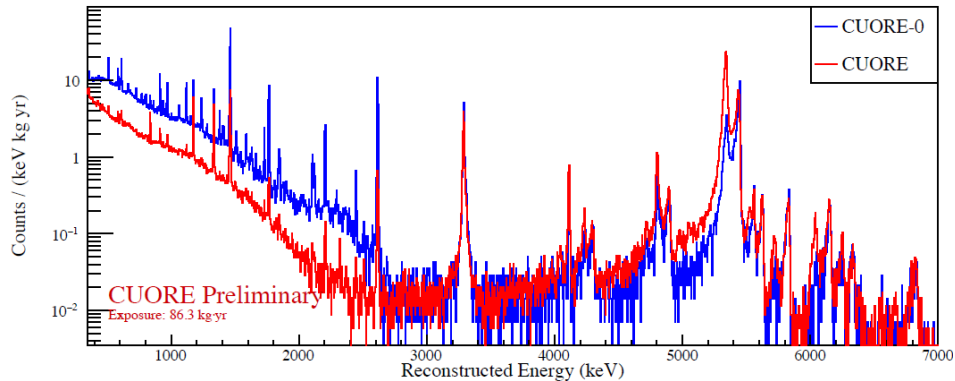


1300 kg of  $^{130}\text{Te}$

See talk on Sept 11  
by Edward Leming



# $0\nu\beta\beta$ -decay with bolometers CUORE and CUPID



Method	Bolometers
Location	LNGS, Italy
Isotope	$^{130}\text{Te}$
$T_{1/2}$ sensitivity	$> 0.7 \cdot 10^{25}$ yr (90% CL)
Limit (latest)	$> 1.5 \cdot 10^{25}$ yr (90% CL)

*PRL 120 132501 (2018)*

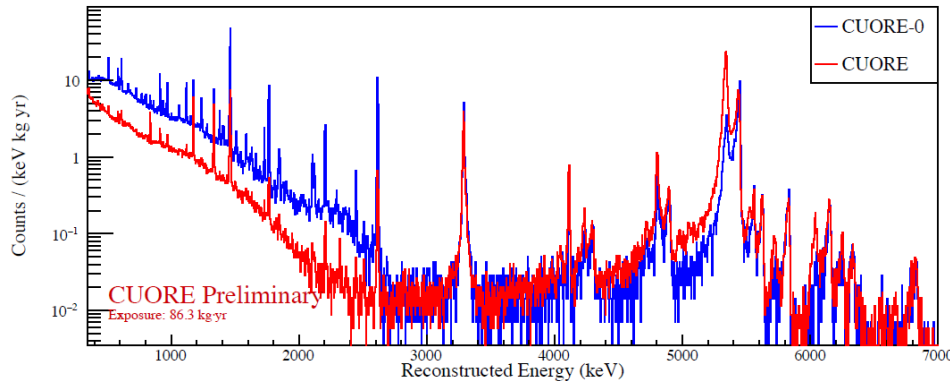
Status 2018:

- ✓ CUORE is taking data
- ✓ 5 y projected half-life sensitivity:  $\sim 10^{26}$  y

See today's talk  
by Paolo Gorla



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Location	LNGS, Italy
Isotope	$^{130}\text{Te}$
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*PRL 120 132501 (2018)*

Status 2018:

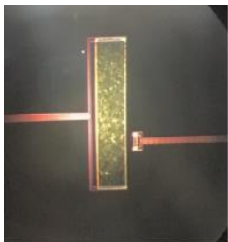
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See today's talk  
by Paolo Gorla

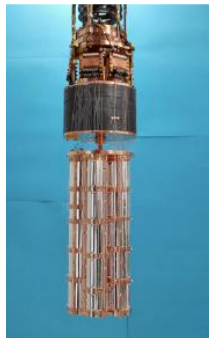
### CUORE Upgrade with Particle ID (CUPID)

New detector technology – luminescent bolometers:

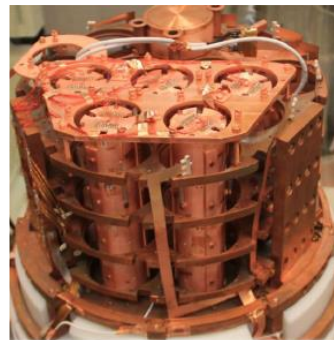
$^{130}\text{TeO}_2$  + Cherenkov light



CUPID-0 –  $\text{Zn}^{82}\text{Se}$



CUPID-Mo –  $\text{Li}_2^{100}\text{MoO}_4$  – baseline option for CUPID

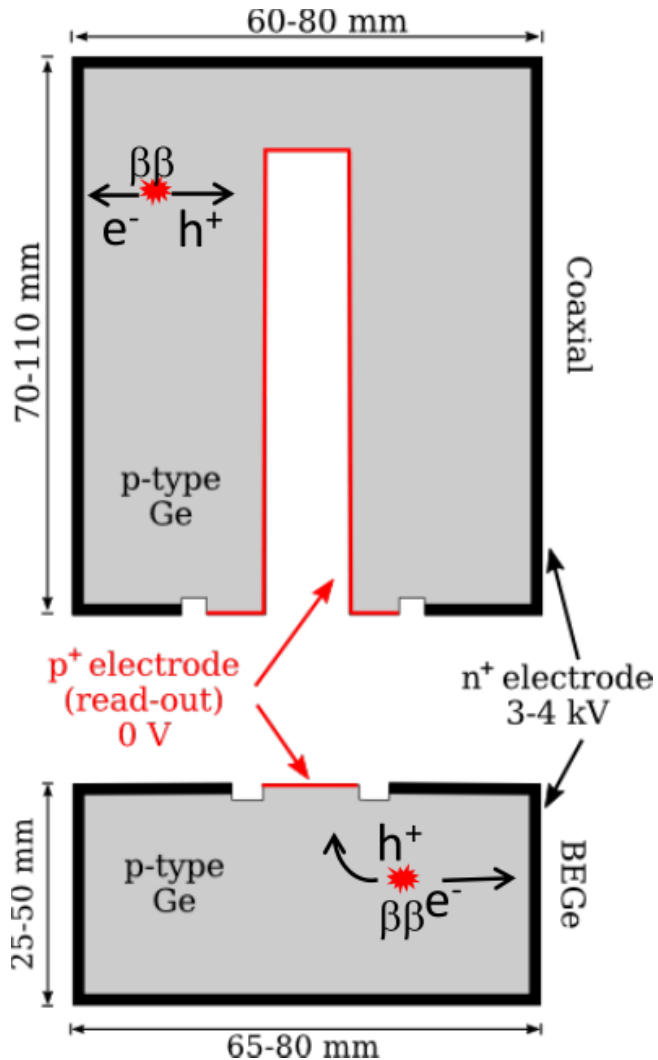


Mission:  
half-life sensitivity higher than  $10^{27}$  y

See today's talk  
by Nicola Casali

# $0\nu\beta\beta$ -decay with Ge detectors

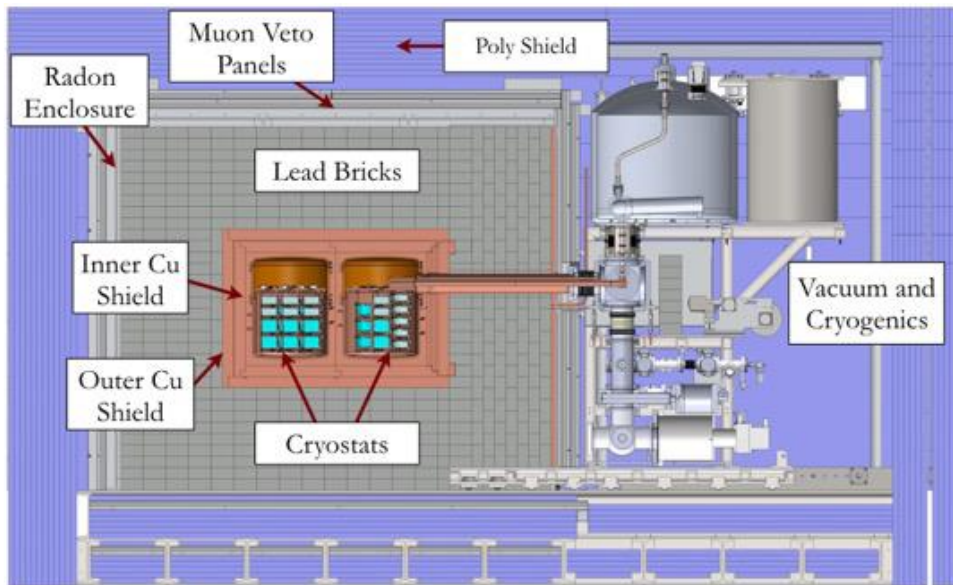
## HPGe detectors enriched in $^{76}\text{Ge}$



- ✓ detector-grade germanium is high-purity material  
⇒ low background
- ✓ established detector technology  
⇒ industrial support
- ✓ very good energy resolution  
~0.1% at  $Q_{\beta\beta}$
- ✓ high detection efficiency  
source = detector

# $0\nu\beta\beta$ -decay with Ge detectors

## MJD



Features:

- ✓ Radiopurity of nearby parts (FETs, cables, Cu mounts, etc.)
- ✓ Low noise electronics yields better PSD
- ✓ Low energy threshold (cosmogenic and low-E background)

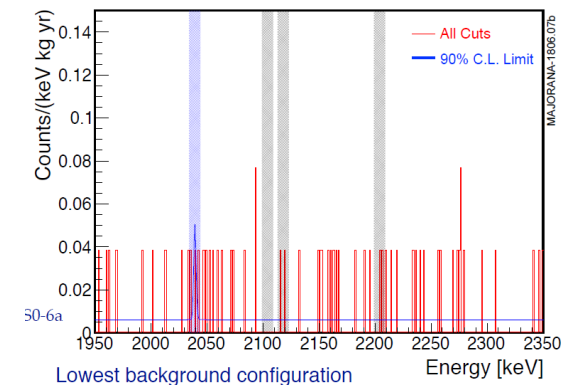
Status 2018:

- ✓ Data taking ongoing
- ✓ Planning an upgrade to improve channel reliability and background

Method	Ge detectors
Location	SURF, USA
Isotope	$^{76}\text{Ge}$
$T_{1/2}$ sensitivity	$> 4.8 \cdot 10^{25}$ yr (90% CL)
Limit (latest)	$> 2.7 \cdot 10^{25}$ yr (90% CL)

*V. Guiseppe, Neutrino2018*

- ✓ Arrays of Ge-diodes in high purity electroformed Cu cryostat

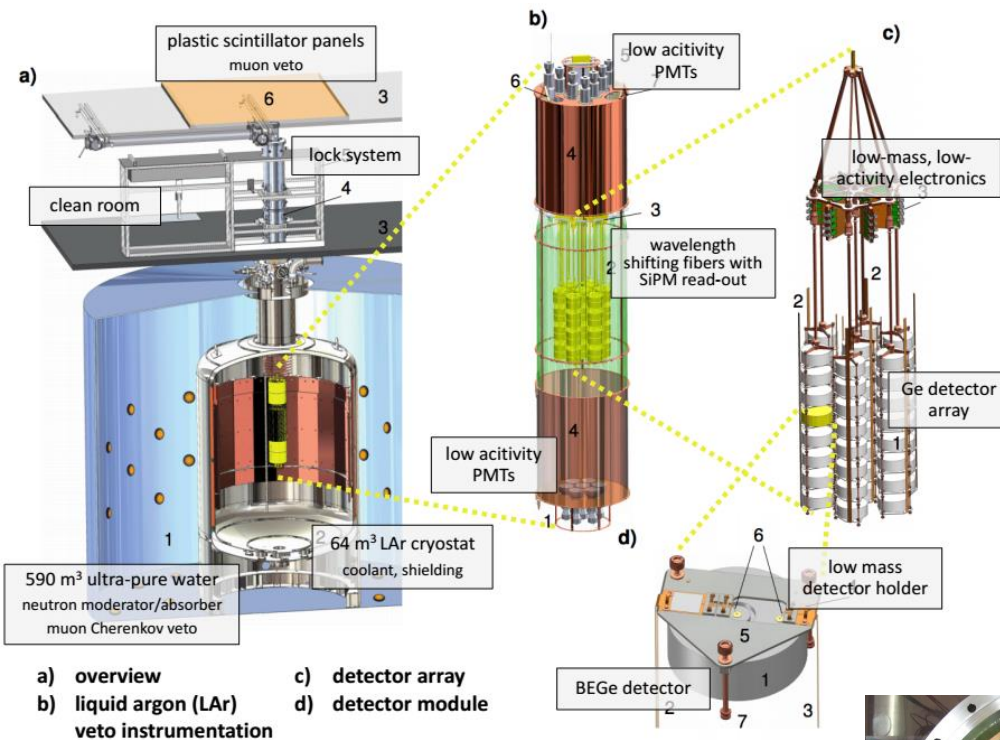


Expect to reach 50-70 kg yr exposure with sensitivity in the range of  $10^{26}$  yr



# $0\nu\beta\beta$ -decay with Ge detectors

## GERDA



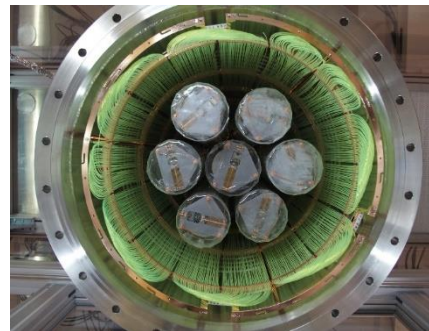
### Features:

- ✓ LAr active veto
- ✓ Low-A shield, no Pb

Method	Ge detectors
Location	LNGS, Italy
Isotope	$^{76}\text{Ge}$
$T_{1/2}$ sensitivity	$> 1.1 \cdot 10^{26}$ yr (90% CL)
Limit (latest)	$> 0.9 \cdot 10^{26}$ yr (90% CL)

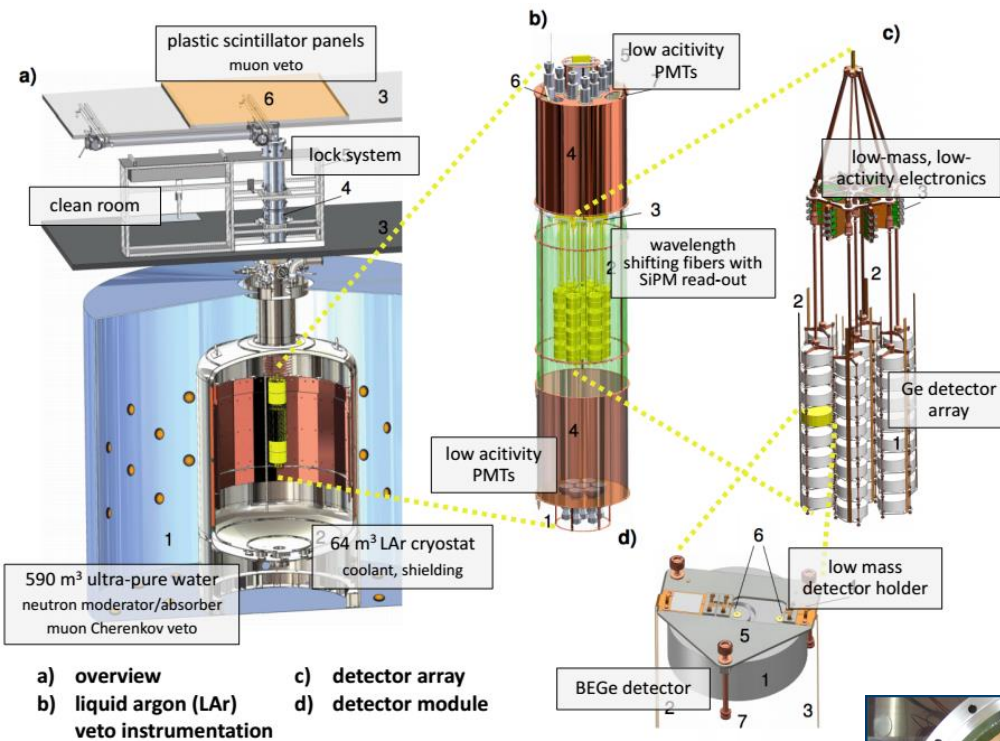
*A.J. Zsigmond, Neutrino2018*

✓ Bare Ge-diodes array in liquid Ar



GERDA will collect data until the end of 2019

# $0\nu\beta\beta$ -decay with Ge detectors GERDA



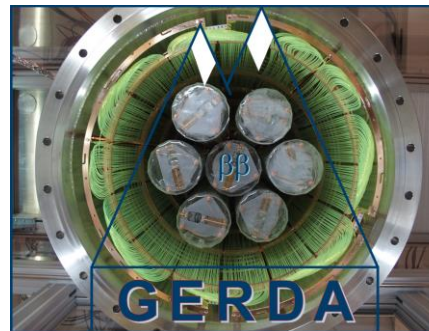
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*A.J. Zsigmond, Neutrino2018*

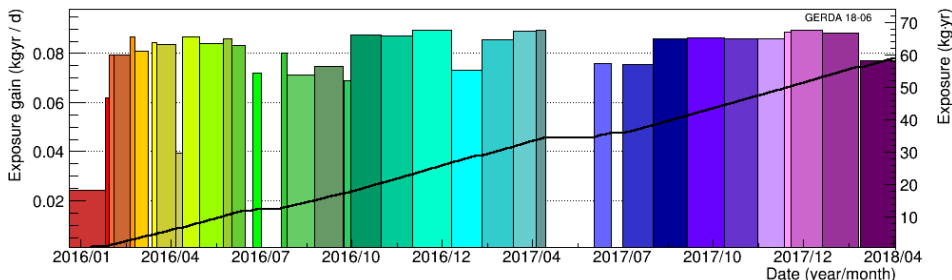
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# $0\nu\beta\beta$ -decay with Ge detectors

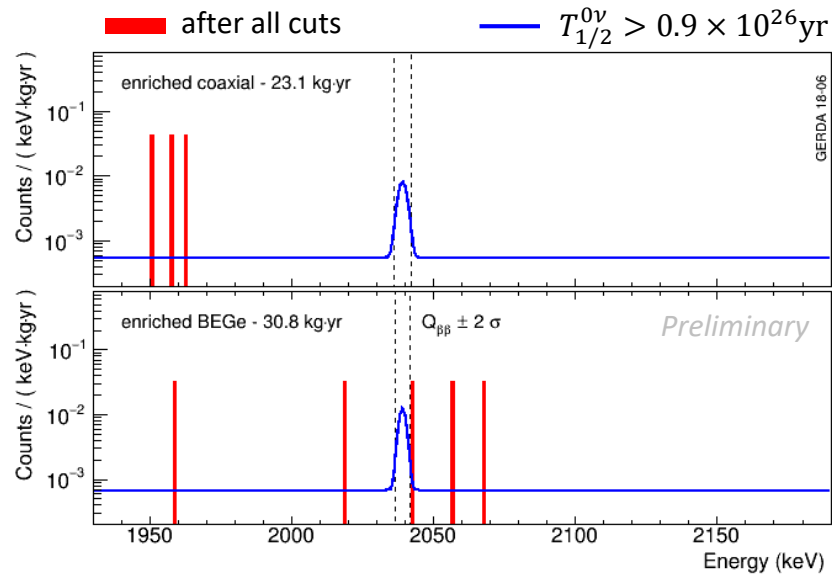
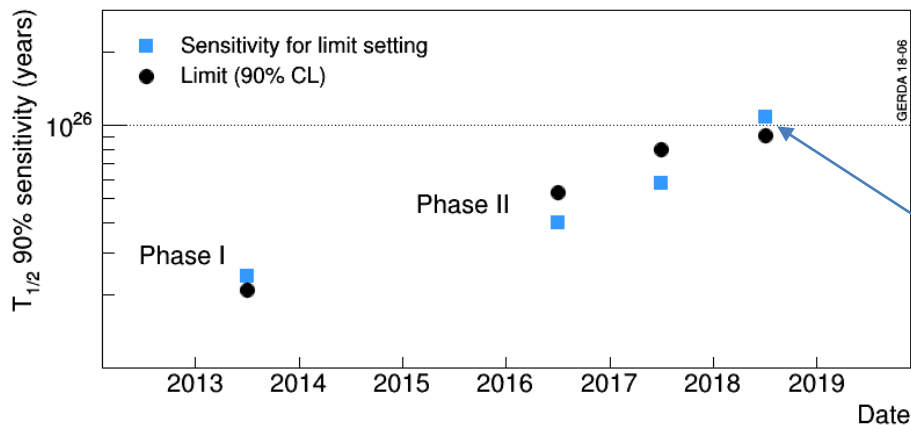
## GERDA: results 2018



- ✓ 60 kg yr of data collected in Phase II by April 2018
  - ✓ 82.4 kg yr in total (Phase I + II)
  - ✓ **Unique** background indices achieved:
    - Coax:  $5.7_{-2.6}^{+4.1} \times 10^{-4}$  cts/(keV·kg·yr)
    - BEGe:  $5.6_{-2.4}^{+3.4} \times 10^{-4}$  cts/(keV·kg·yr)
- best in the field when normalized to FWHM!*

### New 2018 limits:

- ✓ Median sensitivity for limit setting:  
 $1.1 \times 10^{26}$  yr (world best!)
- ✓ Best fit  $\rightarrow$  no signal  
 $T_{1/2}^{0\nu} > 0.9 \times 10^{26}$  yr (90% CL)
- ✓ Probability to have stronger limit 63%



GERDA reached important milestone for  $0\nu\beta\beta$  search!

See today's talk  
by Christoph Wiesinger



# $0\nu\beta\beta$ -decay with Ge detectors

## GERDA: upgrade 2018

Upgrade of the GERDA experiment aims to:

- ✓ Test the novel detectors + increase the mass of  $^{76}\text{Ge}$
- ✓ Show the possibility to improve the background index
- ✓ Prove the robustness and reproducibility of the GERDA approach

Upgrade includes:

- New LAr veto:
  - ✓ new fiber curtain (improved light collection) + central module to read out hidden Ar volume
- Installation of 5 novel **inverted coaxial detectors** made from  $^{76}\text{Ge}$ 
  - ✓ Total increase of  $^{76}\text{Ge}$  mass  $\sim 6$  kg!
- Exchange of all signal and HV cables by new ones with better radiopurity
- New signal cable routing to reduce the cross-talk and improve resolution
- Repairing of broken electronic channels and installation of protective diodes



old curtain



810 fiber ends, 90 SiPMs

new curtain



1215 fiber ends, 135 SiPMs

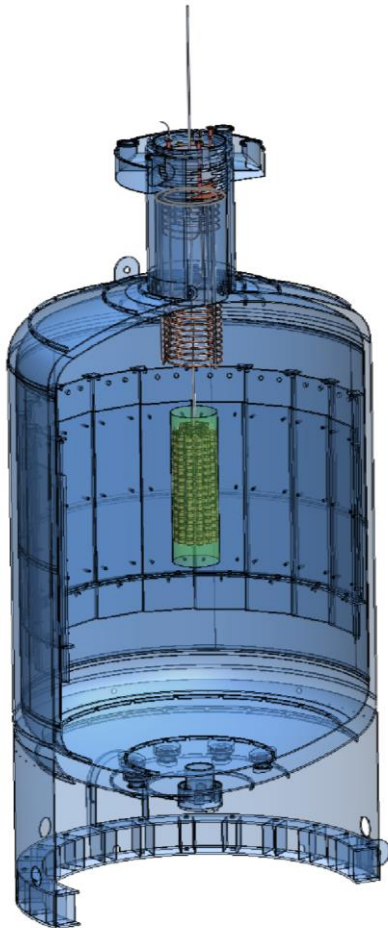
Central module



# $0\nu\beta\beta$ -decay with Ge detectors

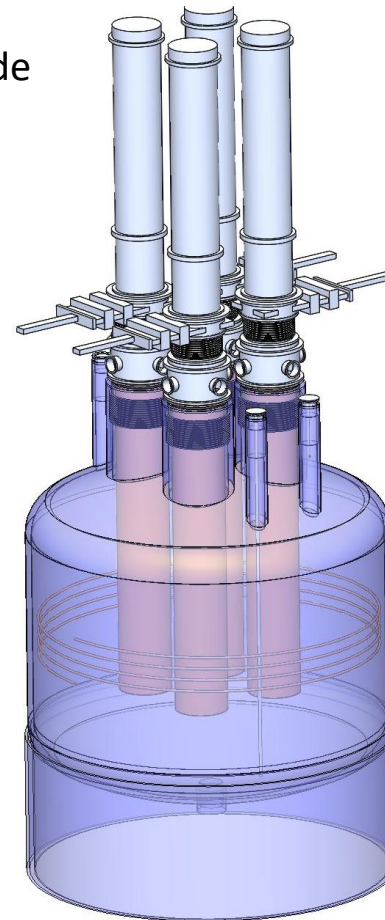
## LEGEND: the best from GERDA and MJD

### Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay



#### First stage (L200):

- ✓ (up to) 200 kg in upgrade of existing GERDA infrastructure at LNGS
- ✓ bkg reduction by factor 3-5 w.r.t GERDA
- ✓ Sensitivity  $10^{27}$  yr



#### Subsequent stages:

- ✓ 1000 kg (staged)
- ✓ timeline connected to DOE down select process
- ✓ bkg factor 30 w.r.t GERDA
- ✓ Location tbd
- ✓ Sensitivity  $10^{28}$  yr

# $0\nu\beta\beta$ -decay with Ge detectors

## LEGEND: the best from GERDA and MJD

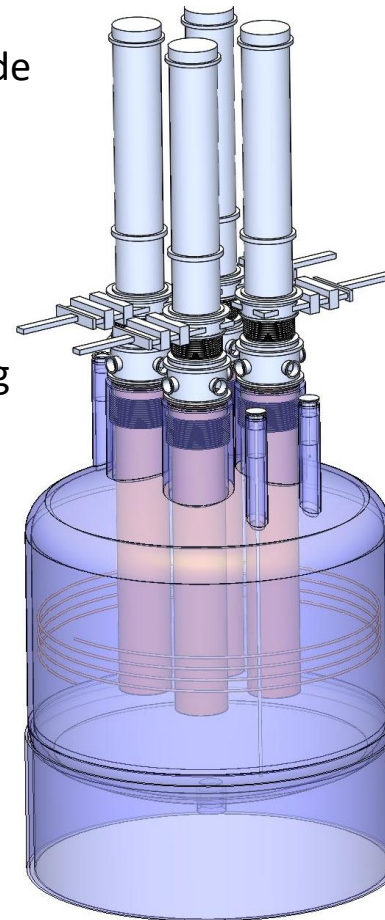
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#### First stage (L200):

- ✓ (up to) 200 kg in upgrade of existing GERDA infrastructure at LNGS
- ✓ bkg reduction by factor 3-5 w.r.t GERDA
- ✓ Sensitivity  $10^{27}$  yr
- ✓ Plan to **start** data taking in **2019**

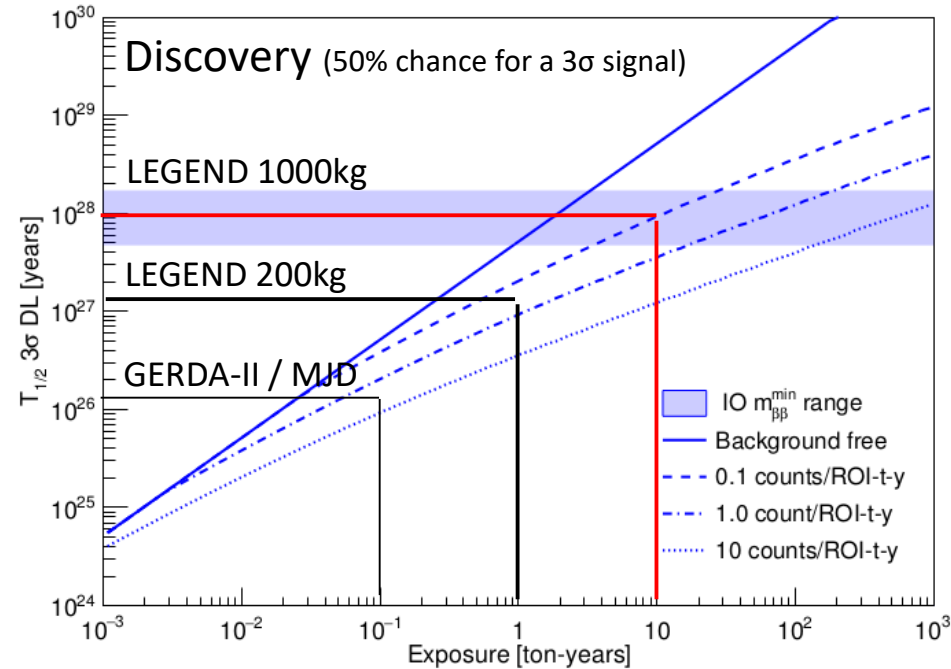
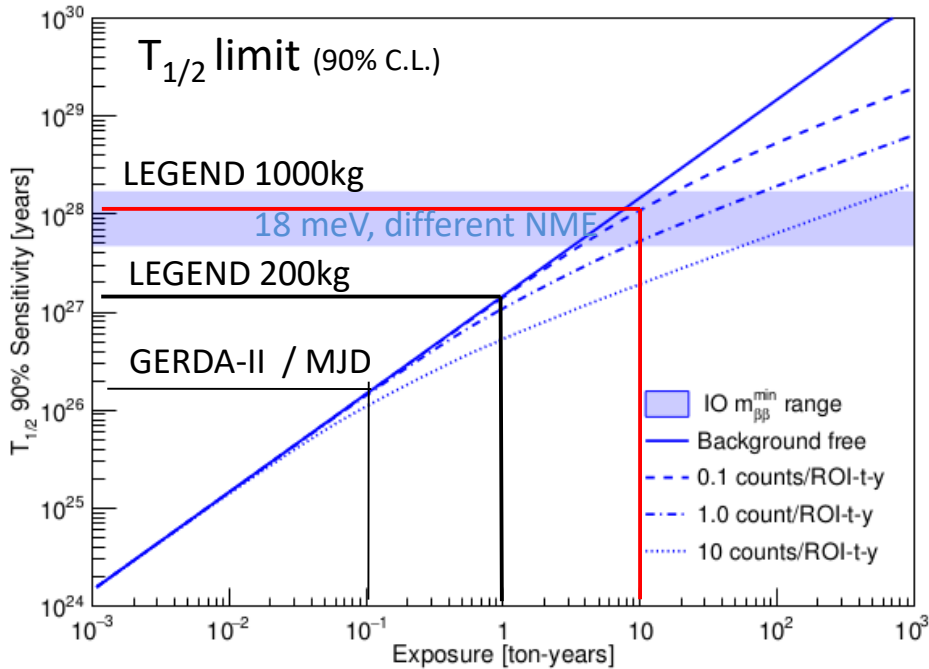
#### Subsequent stages:

- ✓ 1000 kg (staged)
- ✓ timeline connected to DOE down select process
- ✓ bkg factor 30 w.r.t GERDA
- ✓ Location tbd
- ✓ Sensitivity  $10^{28}$  yr



# $0\nu\beta\beta$ -decay with Ge detectors

## LEGEND: sensitivity



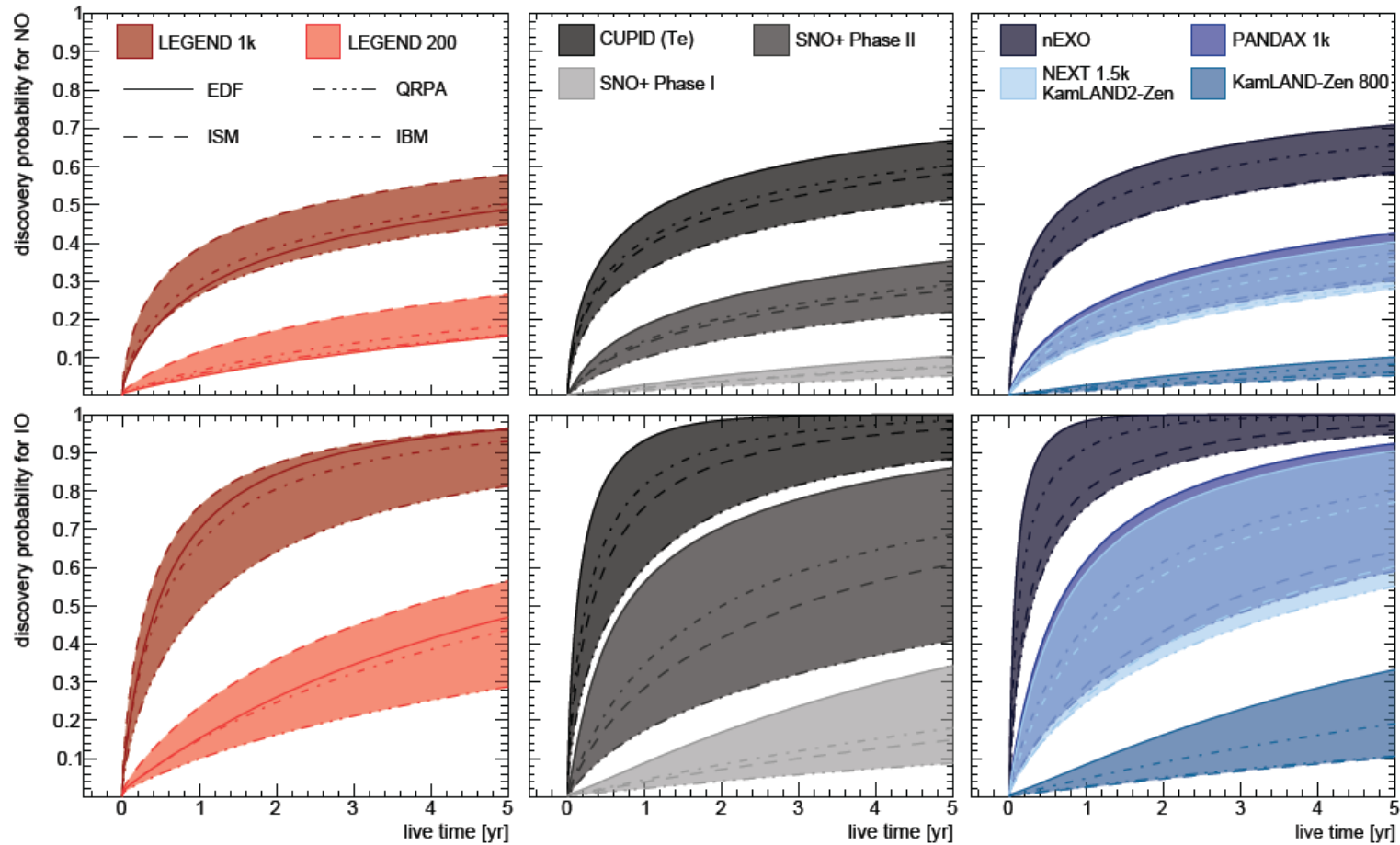
- ✓  $T_{1/2}$  unknown, BSM → 'around corner'
- ✓ background reduction in steps → phased approach

N.B.: background-free operation is a prerequisite for a discovery

- ✓ inputs: 60% efficiency (GERDA number)
- ✓ Background: GERDA/MJD ~ 3 cts/(FWHM t yr)
  - 200 kg ~ 0.6 cts/(FWHM t yr)
  - 1000 kg ~ 0.1 cts/(FWHM t yr)

# $0\nu\beta\beta$ -decay experiments

## Discovery probability



*Discovery probability of next-generation neutrinoless double- $\beta$  decay experiments*  
*M. Agostini, G. Benato and J. A. Detwiler Phys. Rev. D 96, 053001 (2017)*



# $0\nu\beta\beta$ -decay experiments

## Summary

- ✓  $0\nu\beta\beta$  decay is a crucial process, New Physics maybe around the corner
- ✓ Very active field: several ton-scale experiments are in preparation
- ✓ Huge experimental effort: tons of material, but “zero” background
- ✓ The discovery probability for the next generation projects is pretty high
- ✓ We need to observe the signal with multiple isotopes using various experimental methods